

HORIBA

Explore the future

HORIBA Scientific

Raman Spectroscopy

Dr. Christoph Lenz & Dr. Bernd Bleisteiner

Raman Basics & Technique

September 2020

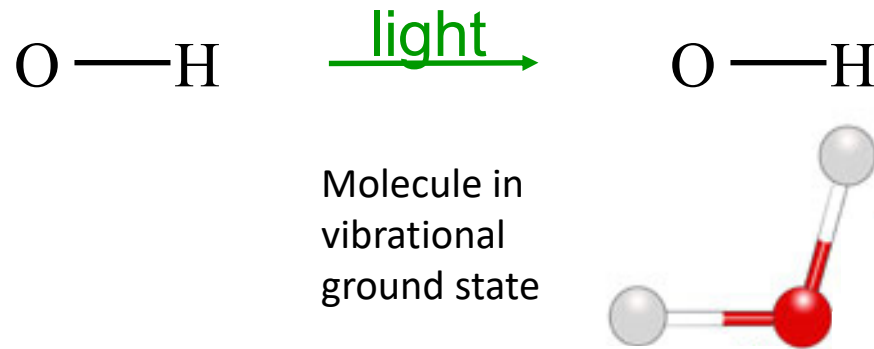


Outline

- Principles of Raman effect in a nutshell
- What information does Raman give?
- Instrument Parameters I
 - Laser
 - Filter / Attenuator
 - Collecting lens
 - Rayleigh Filter
 - Confocal Aperture
 - Slit
 - Spectrometer & Grating
 - Detector & Quantum Efficiency
 - Correction of Instrument Function



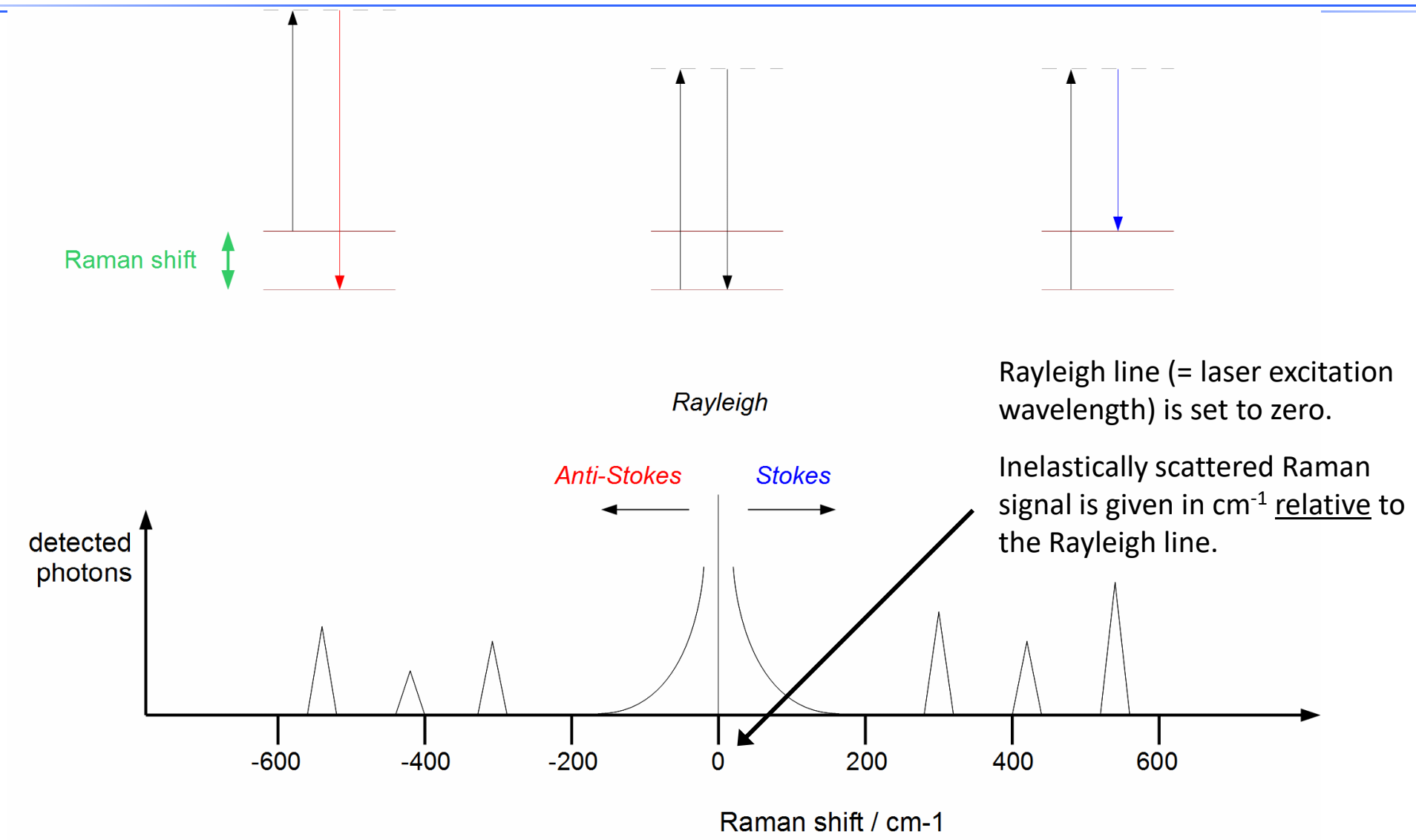
Stokes Raman Effect



Light interacting with molecules causing several effects at the same time:

- 1) There are molecules which remain unchanged → Rayleigh Scattering (elastic scattering)
- 2) There are molecules which are promoted to a higher vibrational energy state → Raman scattering (inelastic scattering). The energy required to do this is taken from the incident light. The light loses energy changing wavelength / colour → Stokes Raman effect

Raman Spectrum



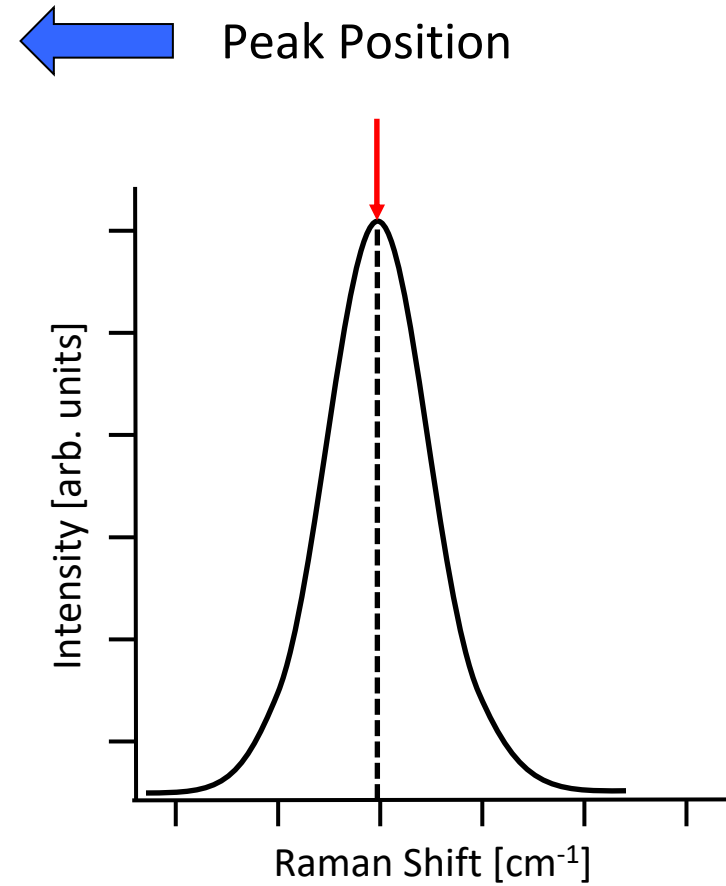
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Raman Information

- Chemical Identity
- Crystal Phases (Polymorphism)

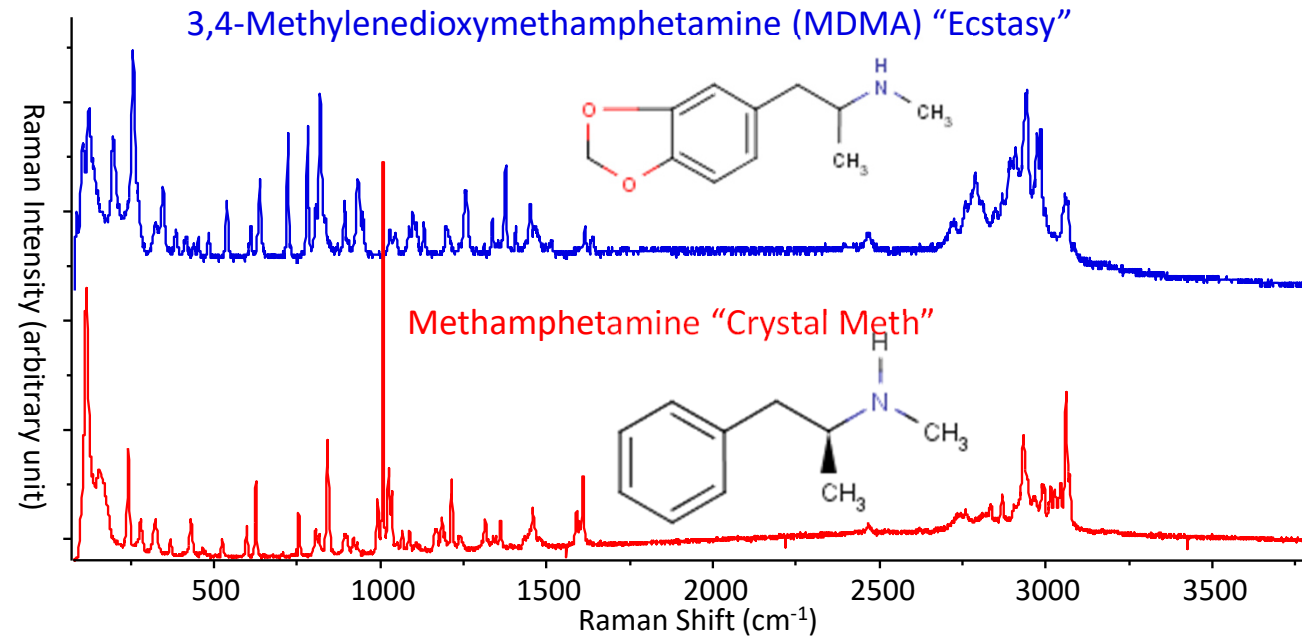


Raman Information

- Chemical Identity
- Crystal Phases (Polymorphism)

← Peak Position

Chemical Identification of Similar Structures

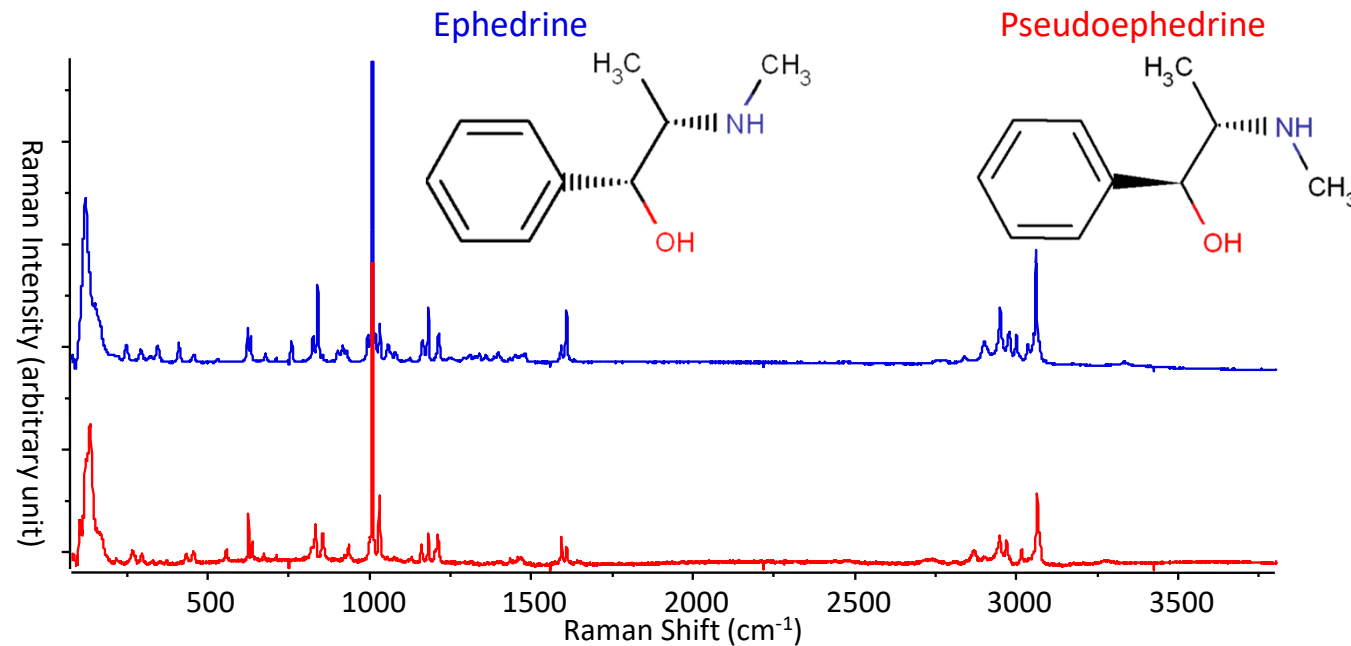


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← Peak Position

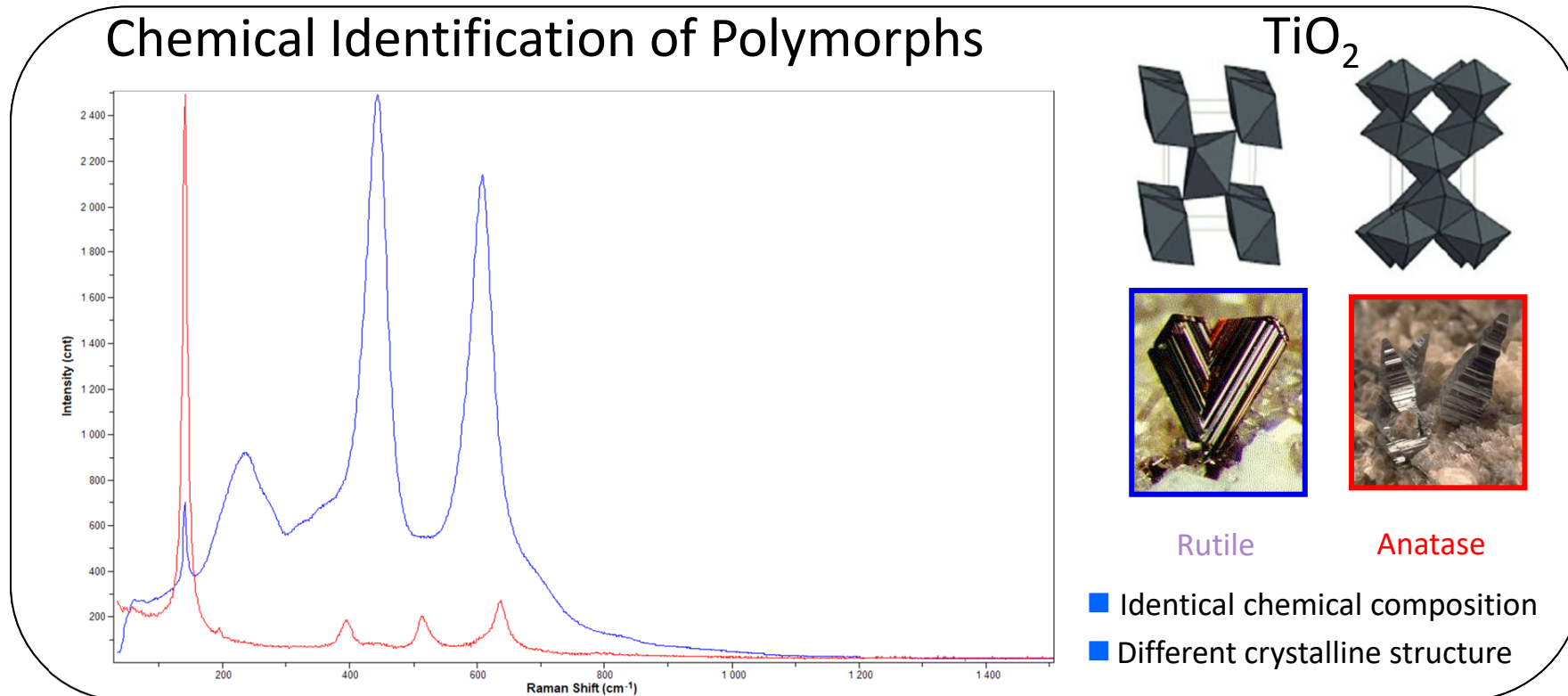
Chemical Identification of Diastereomers



Raman Information

- Chemical Identity
- Crystal Phases (Polymorphism)

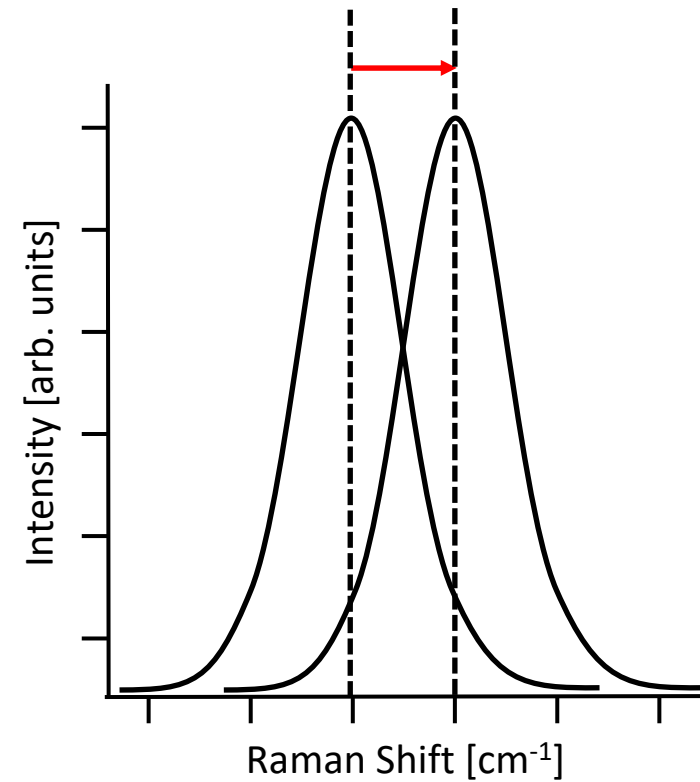
← Peak Position



Raman Information

- Strain, Stresses
- Temperature

← Peak Shift

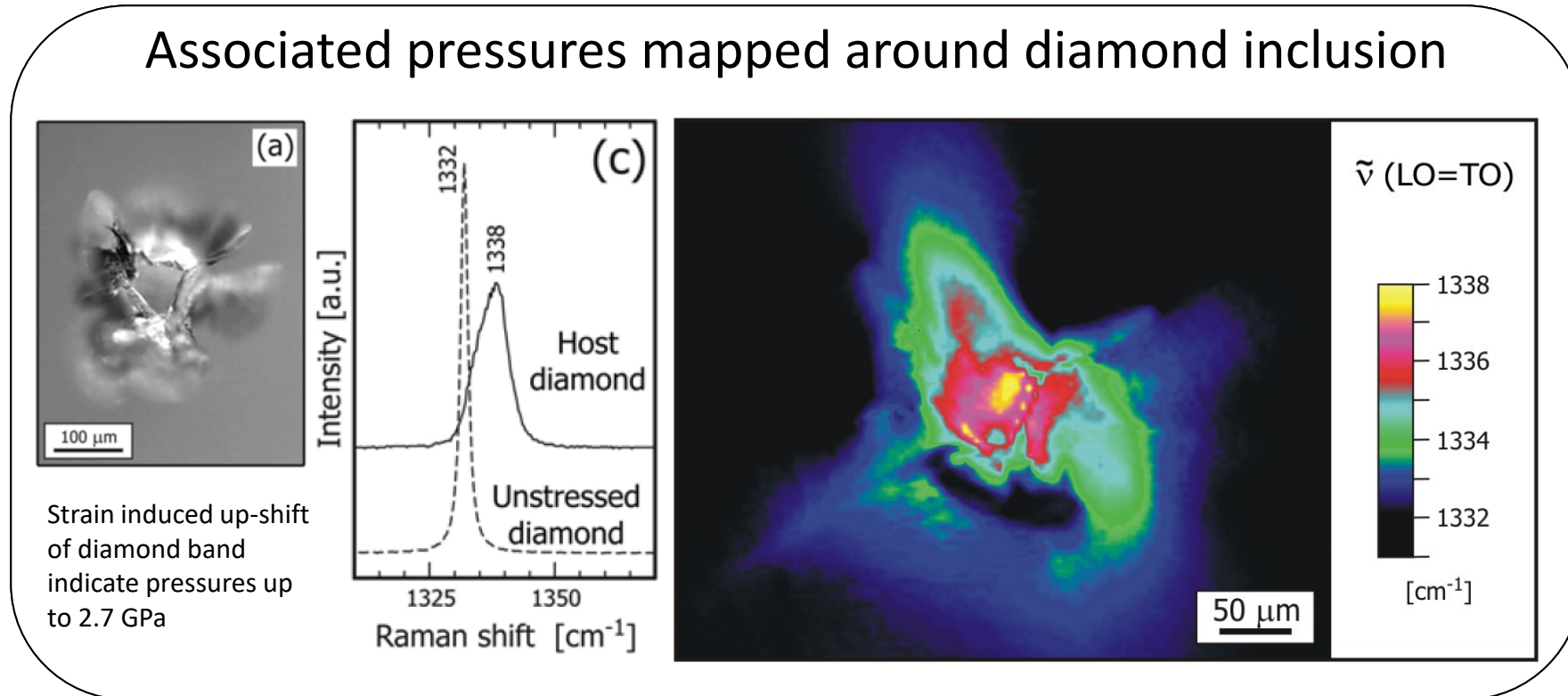


Raman Information

- Strain, Stresses
- Temperature

← Peak Shift

Associated pressures mapped around diamond inclusion

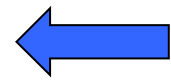


Nasdala, L., Harris, J.W., Hofmeister, W., Micro-spectroscopy of diamond. Asia Oceania Geosciences Society, 4th Annual Meeting, 2007.

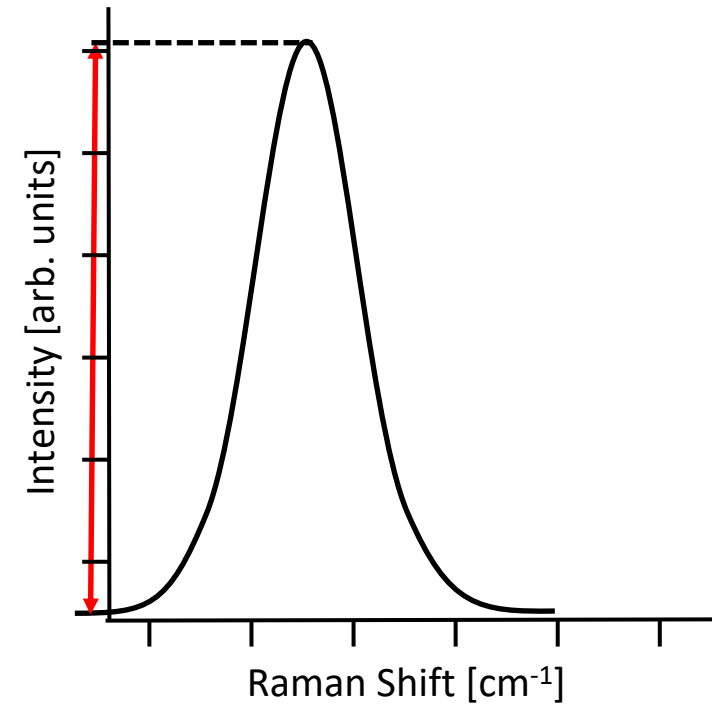
Nasdala, L., Raman barometry of mineral inclusions in diamond crystals. Mitt. Österr. Miner. Ges. 149 (2004)

Raman Information

■ Concentration



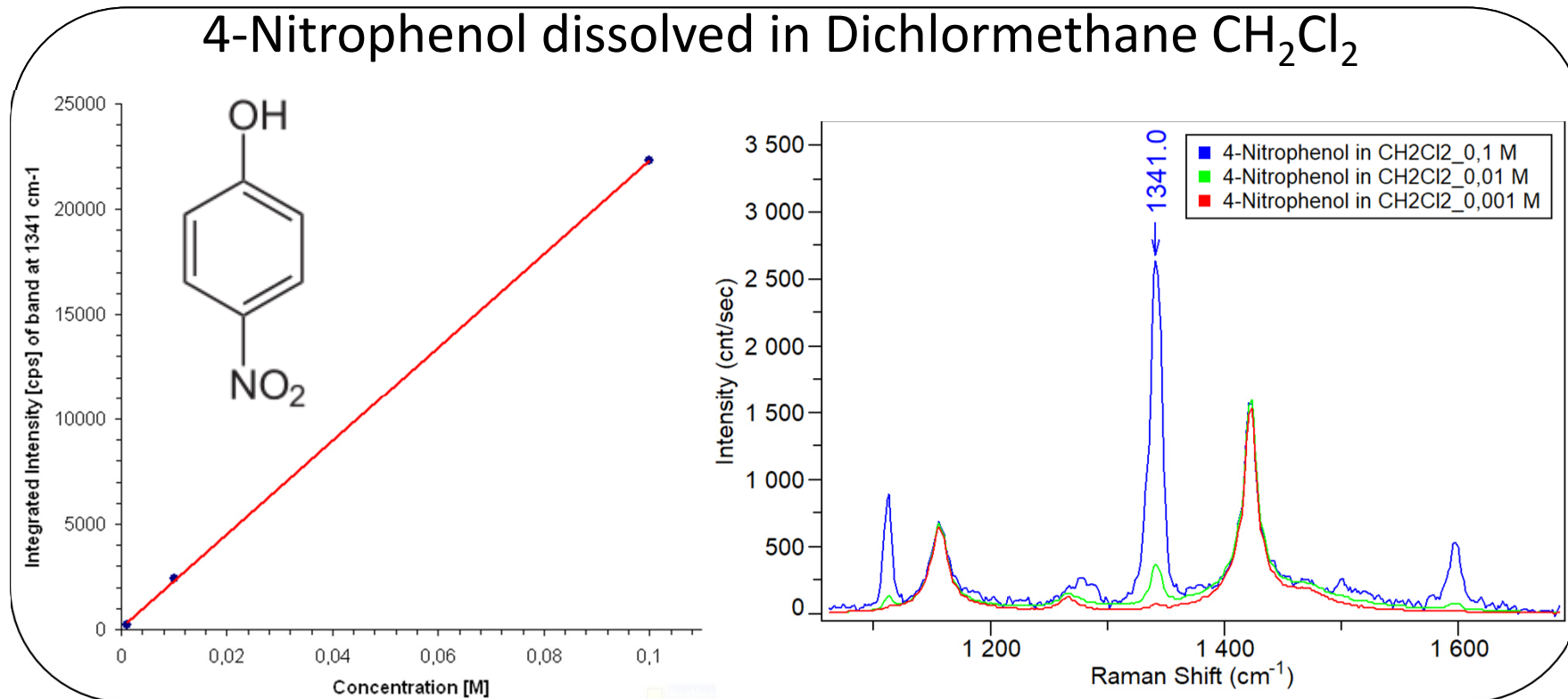
Peak Intensity



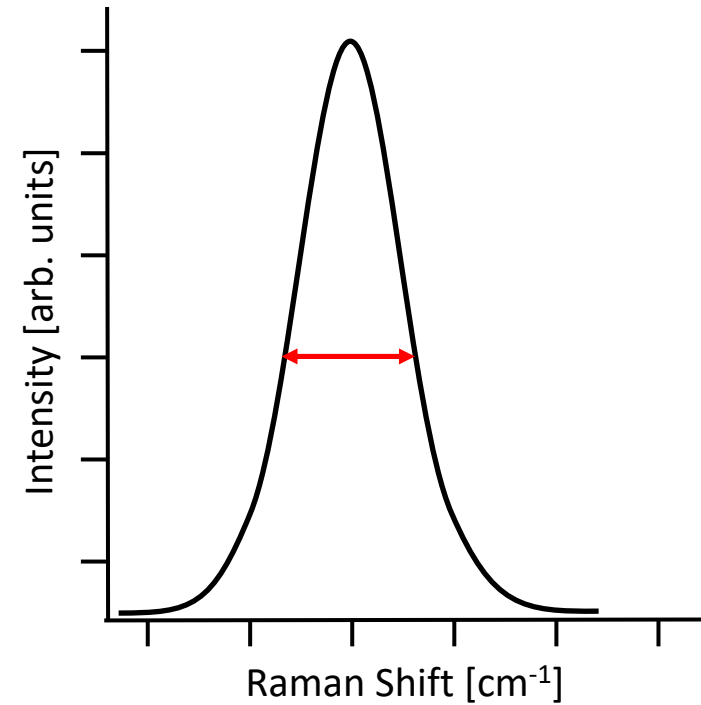
Raman Information

■ Concentration

← Peak Intensity



Raman Information



- Structural Order/Disorder
- Crystallinity

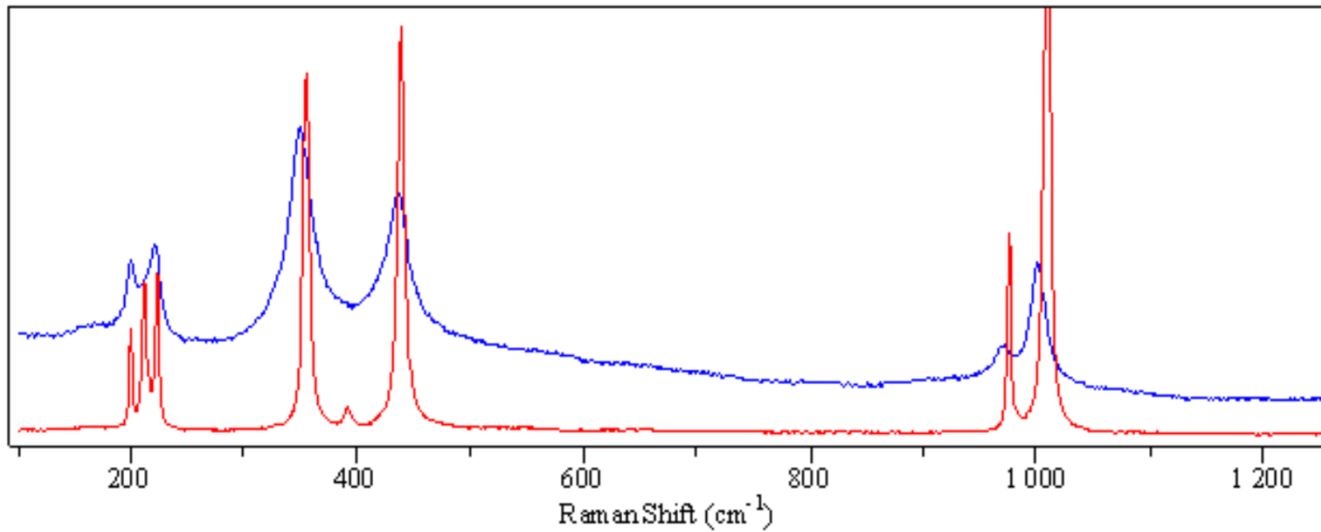
← Peak Width

Raman Information

- Structural Order / Disorder
- Crystallinity

← Peak Width

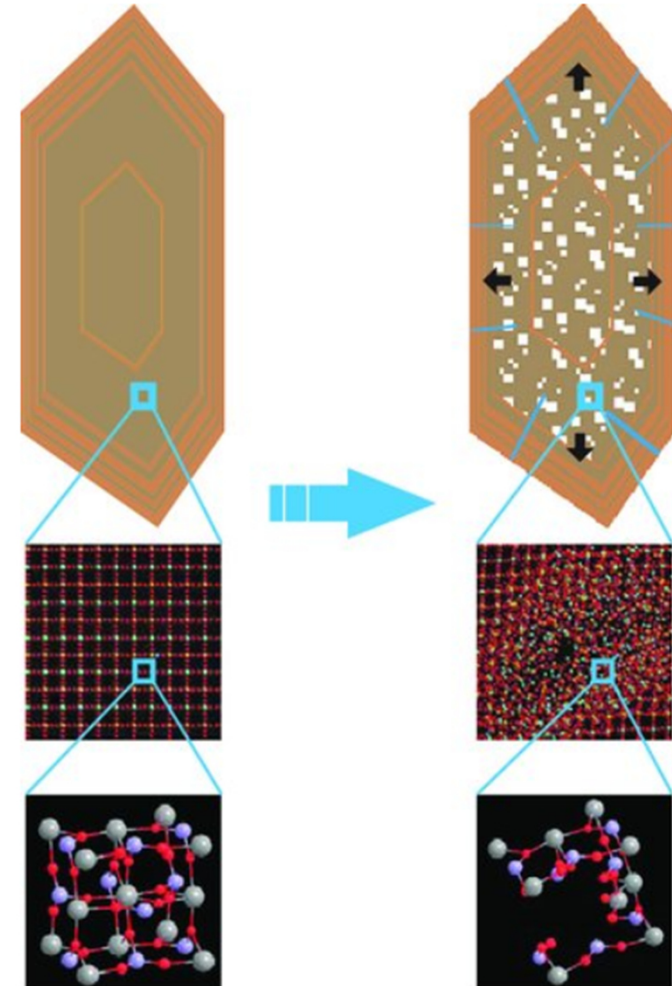
Spectra of crystalline and radiation-damaged zircon



Radiation-damaged zircon (blue)
crystalline zircon (red)

crystalline

Radiation-damaged



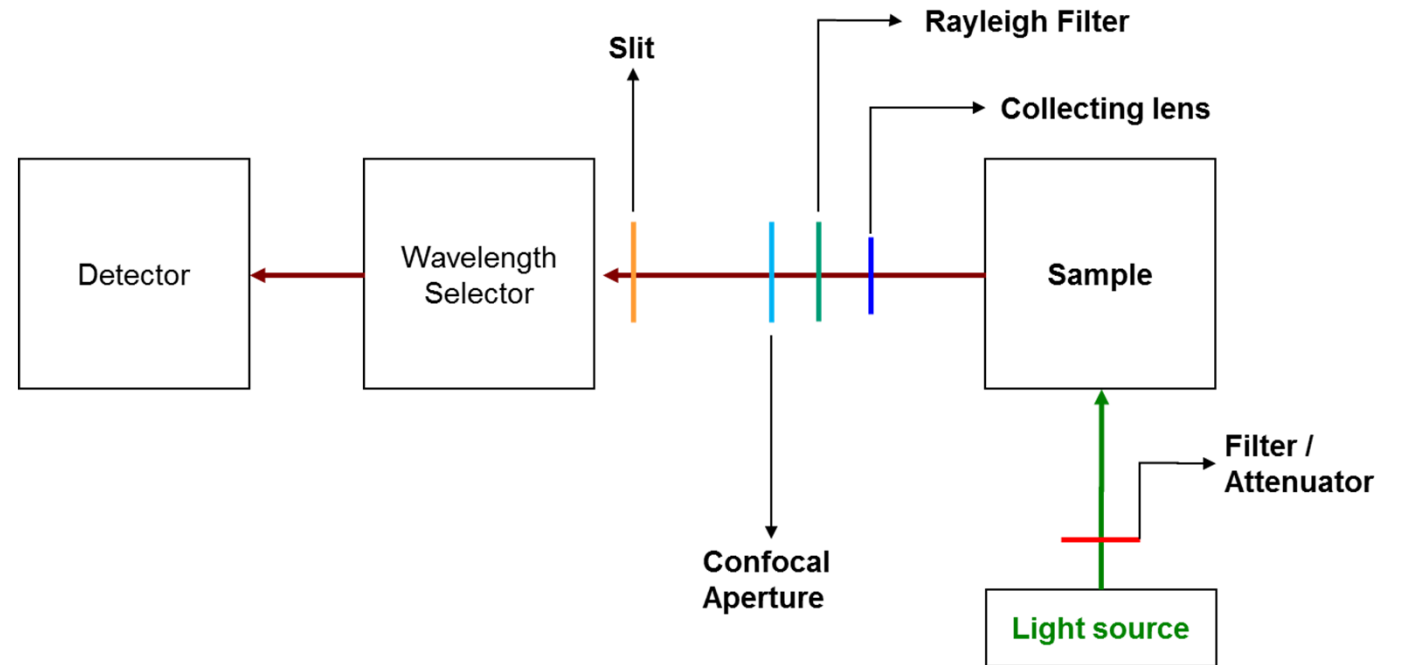
Xu et al. (2012): Reverse age zonation of zircon formed by metamictisation and hydrothermal fluid leaching; Lithos 150:256-267

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- Principles of Raman effect in a nutshell
- What information does Raman give?

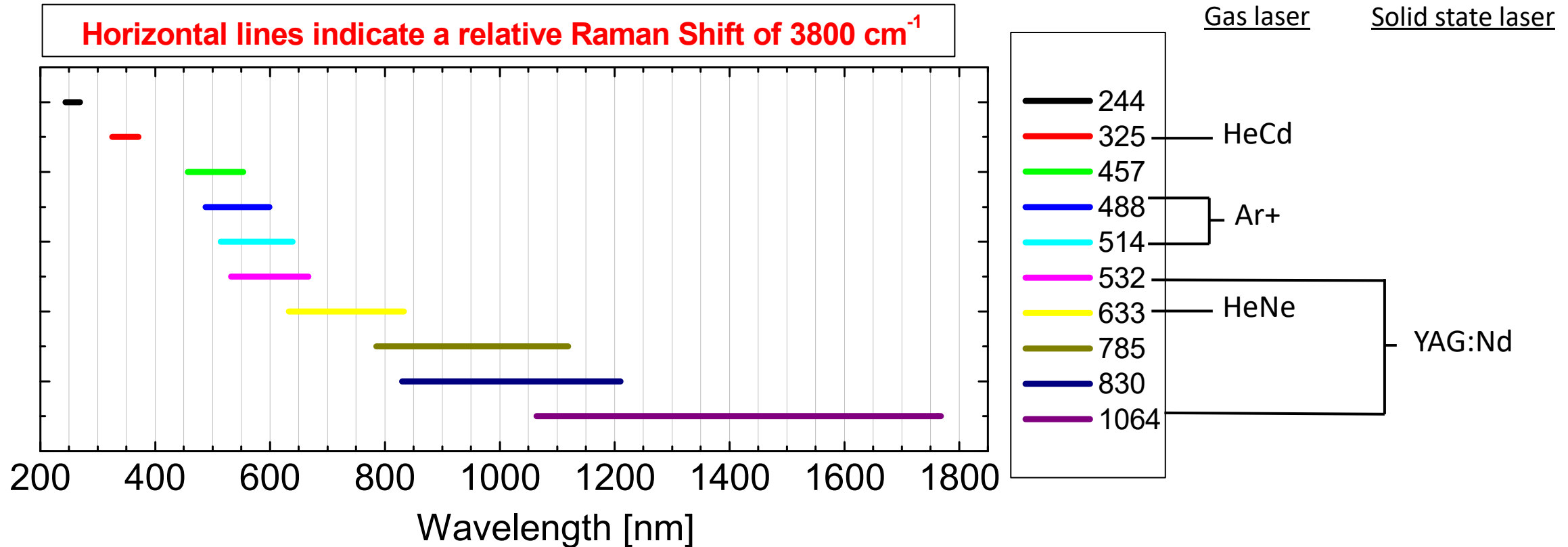
■ Instrument Parameters I

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Laser source

Laser 670 nm

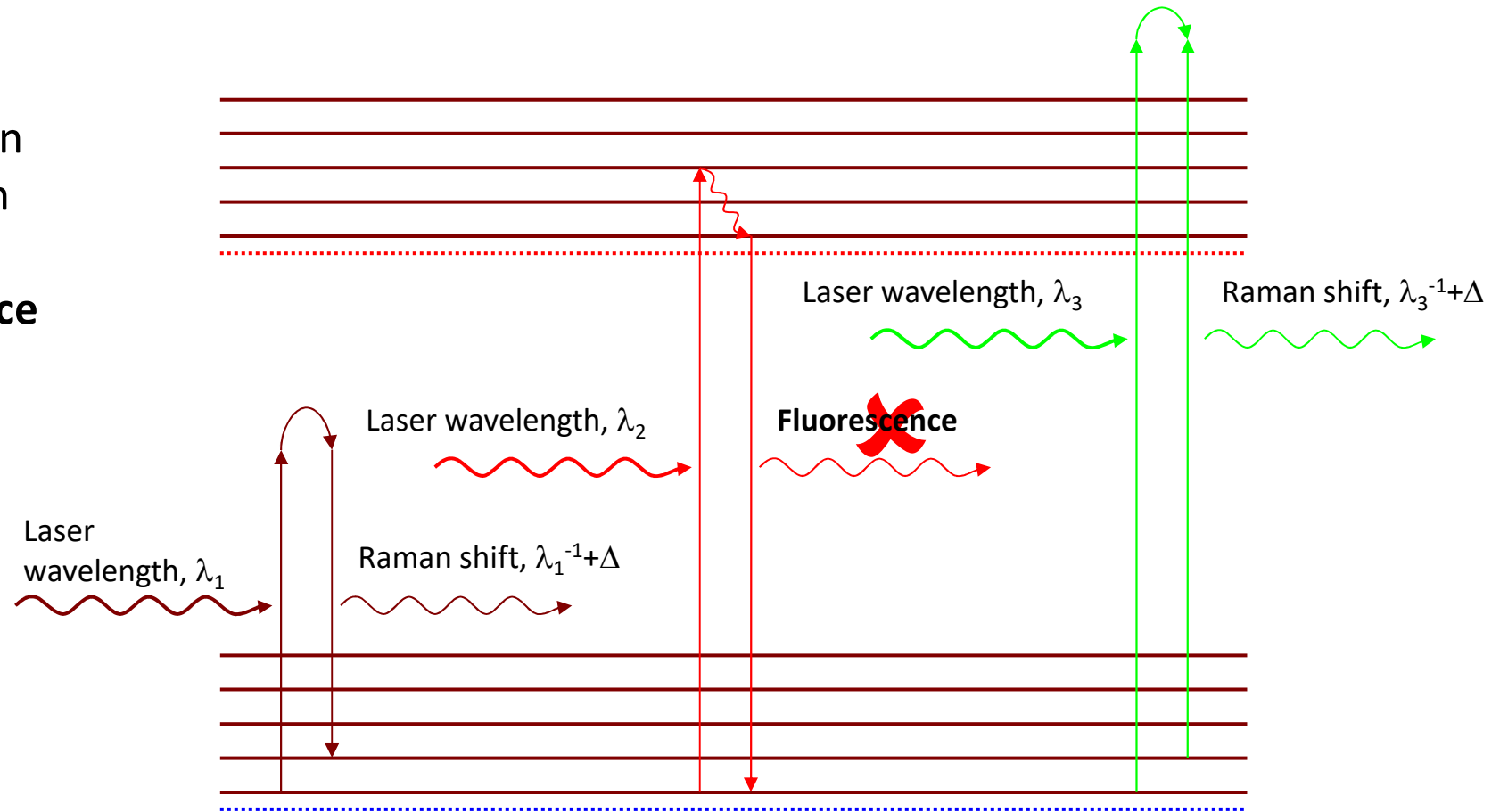


Possible wavelengths are 227, 244, 257, 325, 364, 413, 442, 457, 473, 488, 514, 532, 633, 647, 660, 785, 830, 1064 nm.

Laser source

Laser 670 nm

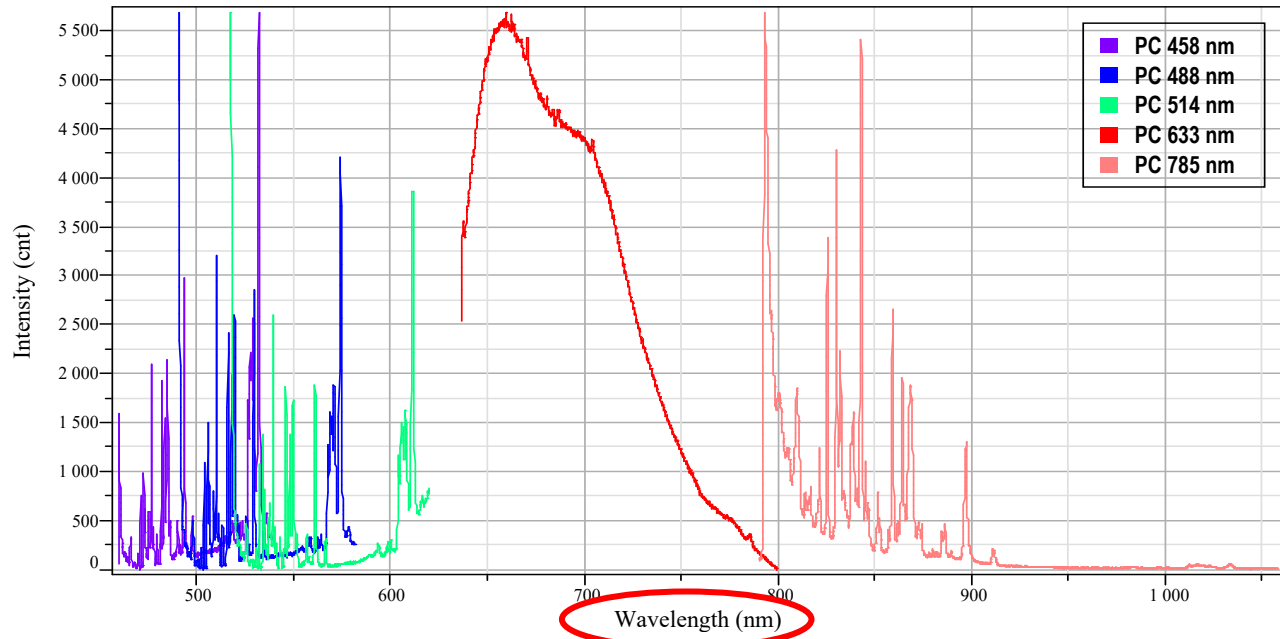
The most important reason for changing the excitation wavelength is to **avoid luminescence/fluorescence**



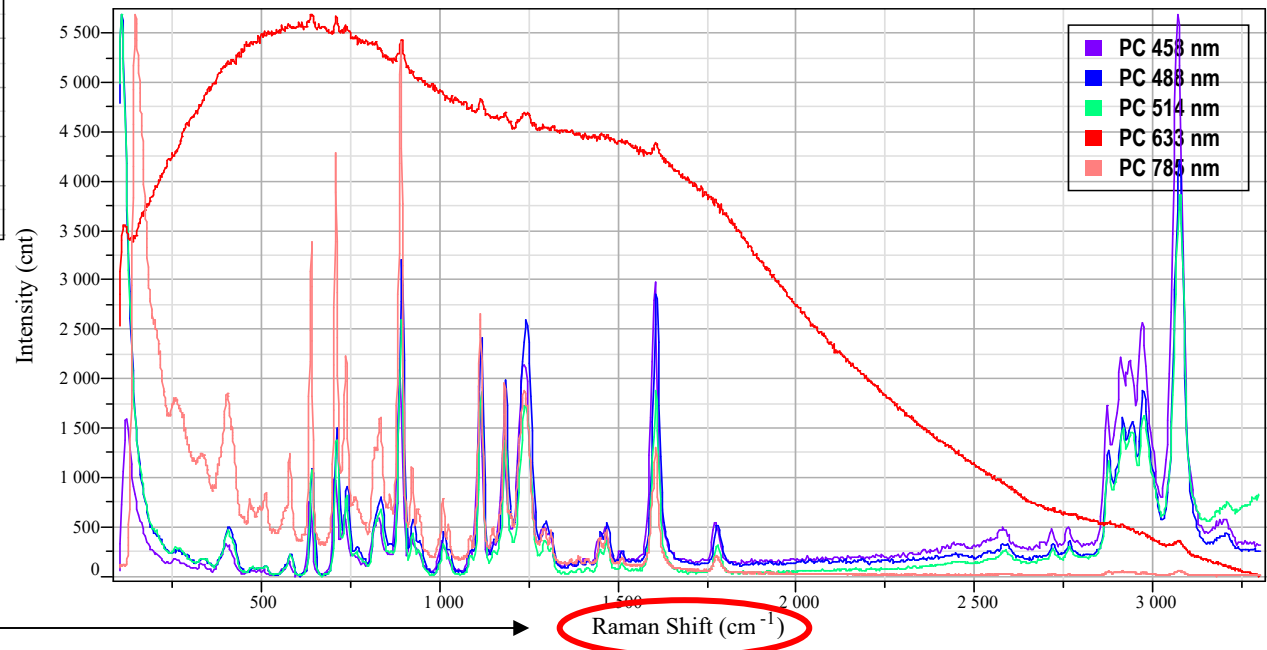
Laser wavelength: $\lambda_1 < \lambda_2 < \lambda_3$

Laser source

Laser 670 nm



Polycarbonat spectra at different excitation wavelengths



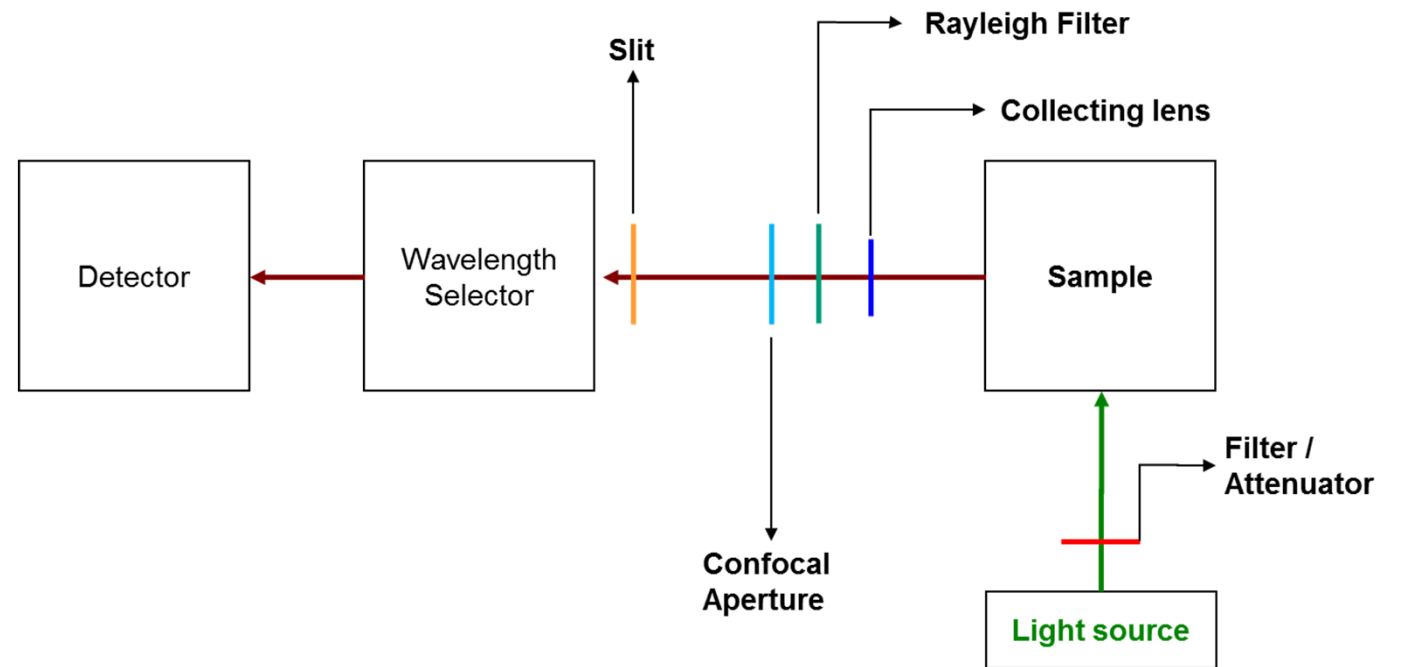
Fluorescence is wavelength dependent
 Ordinary Raman is wavelength independent

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Filter

ND Filter 100%

- Keep in mind: the usage of high numerical objective lenses causes a very small spot size of the laser which results in a **high power density**
- To avoid sample burning radiation power has to be adapted INDIVIDUALLY to the sample

Laser wavelength: 473 nm
Laser power at sample: 25.5 mW

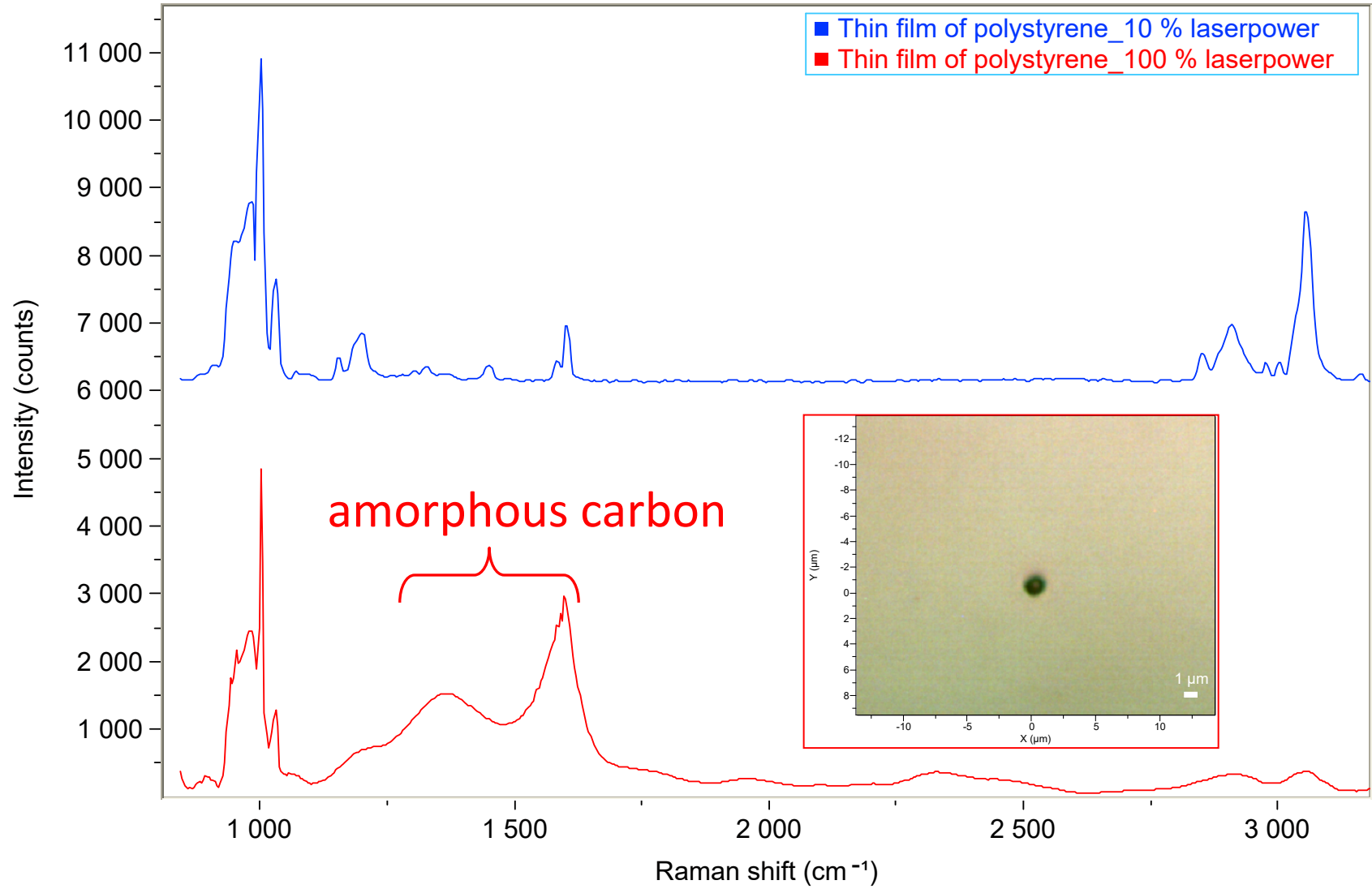
Objective	N.A.	Laser spot size (μm)	Radiation power (kW/cm ²)
100×	0.90	0.64	~7900
50×	0.75	0.77	~5400
10×	0.25	2.31	~600

Laser wavelength: 633 nm
Laser power at sample: 12.6 mW

Objective	N.A.	Laser spot size (μm)	Radiation power (kW/cm ²)
100×	0.90	0.85	~2200
50×	0.75	1.03	~1500
10×	0.25	3.09	~200

Filter

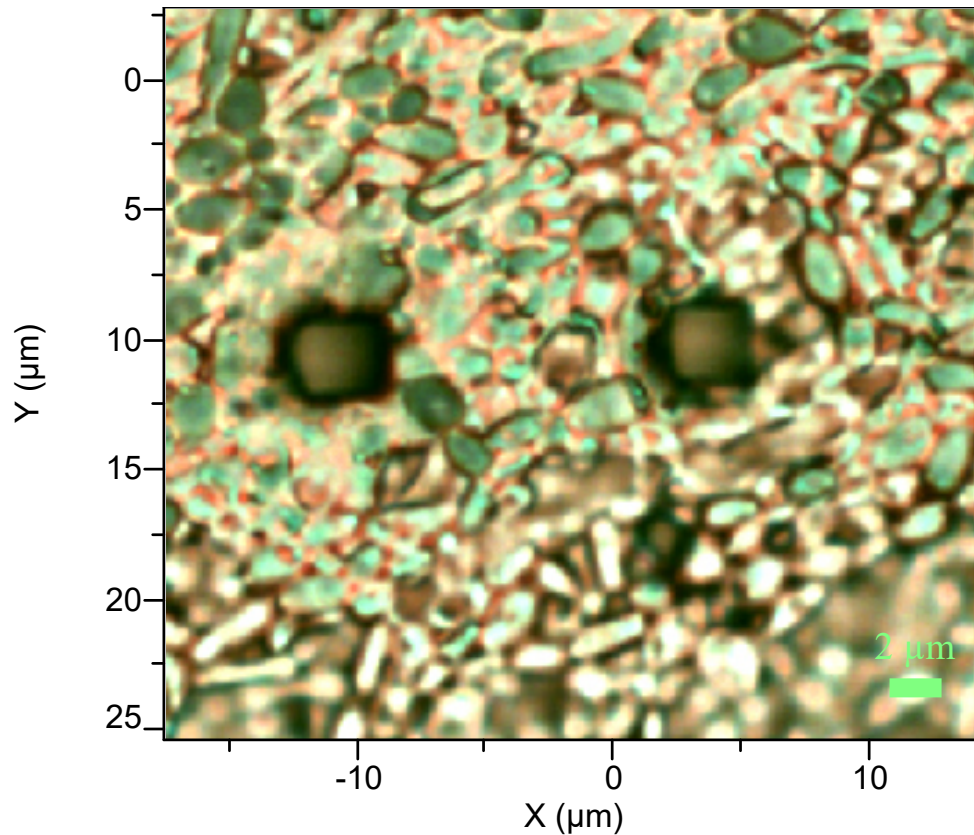
ND Filter 100%



Filter

ND Filter 100%

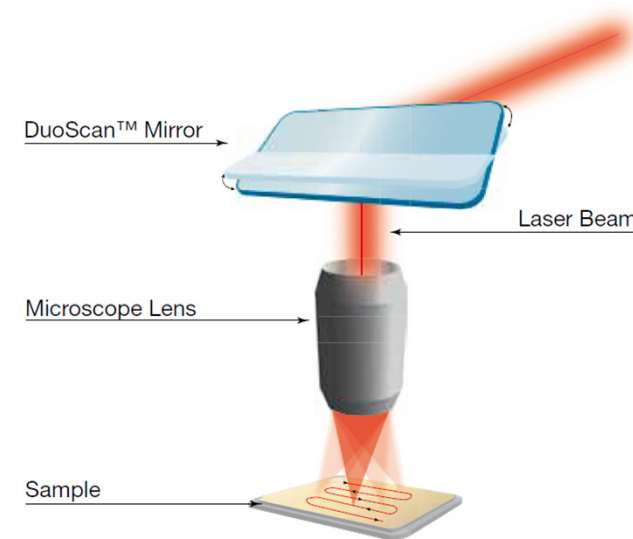
Candida Tropicalis after laser radiation with too much power



Excitation: 473 nm

Laser power > 1 % (0.16 mW)

DuoScan

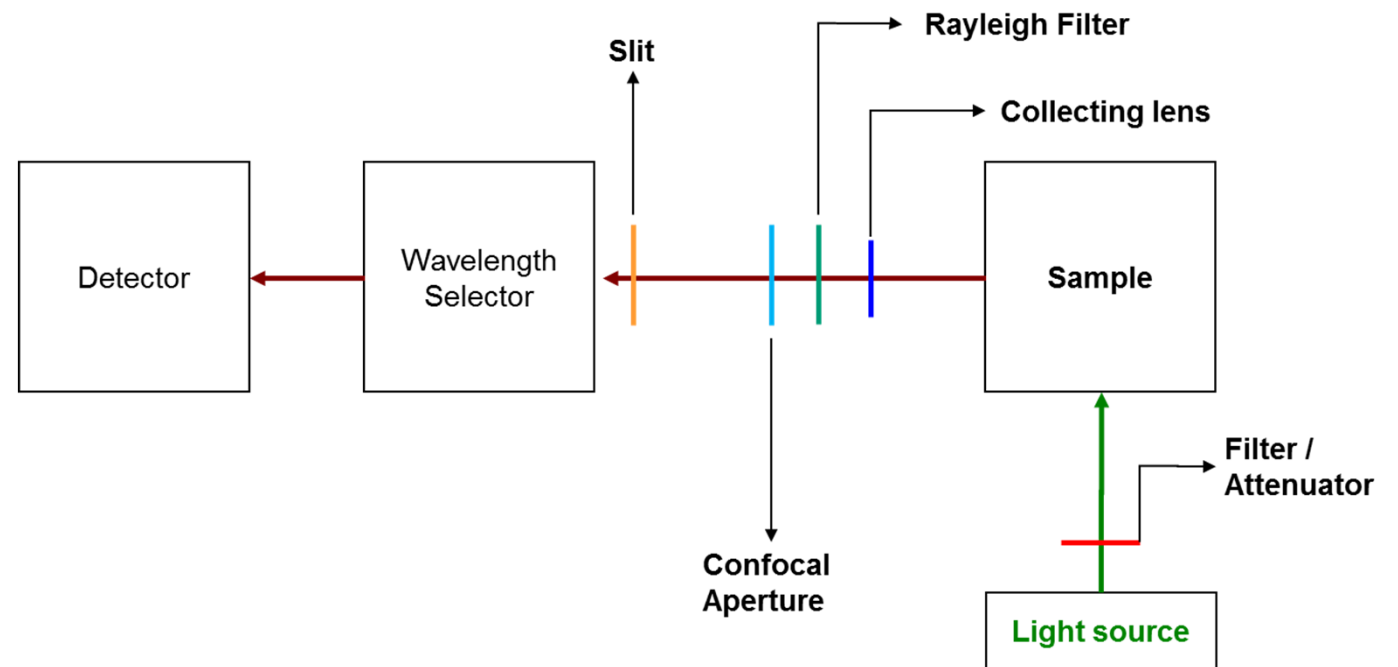


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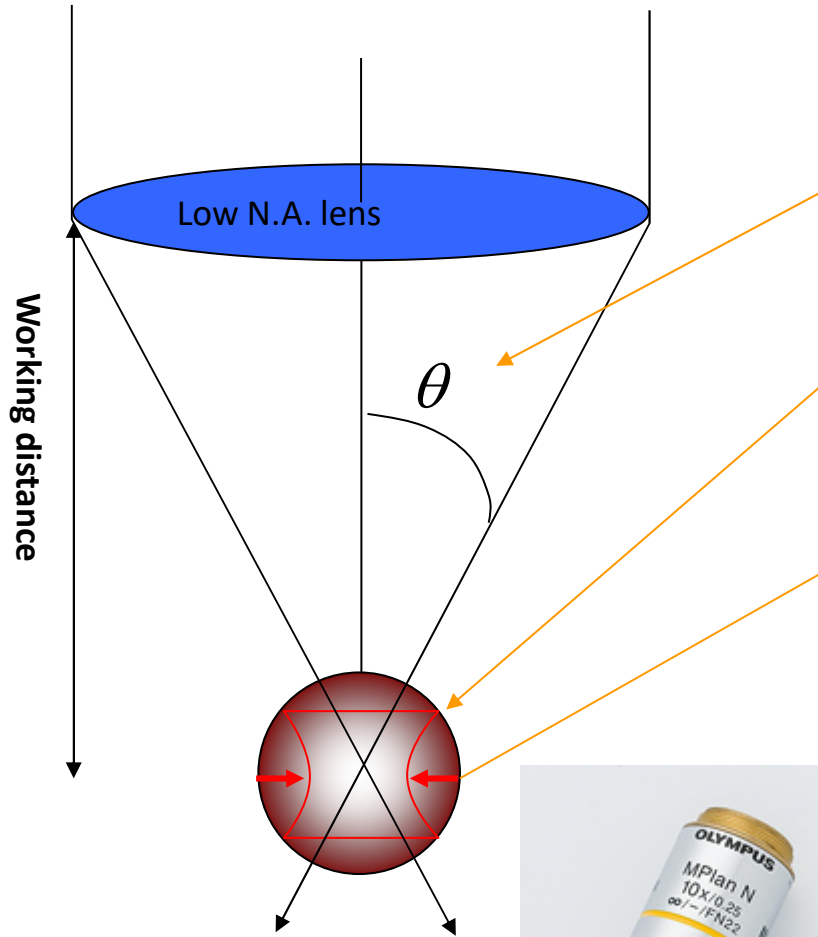
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Collecting lens

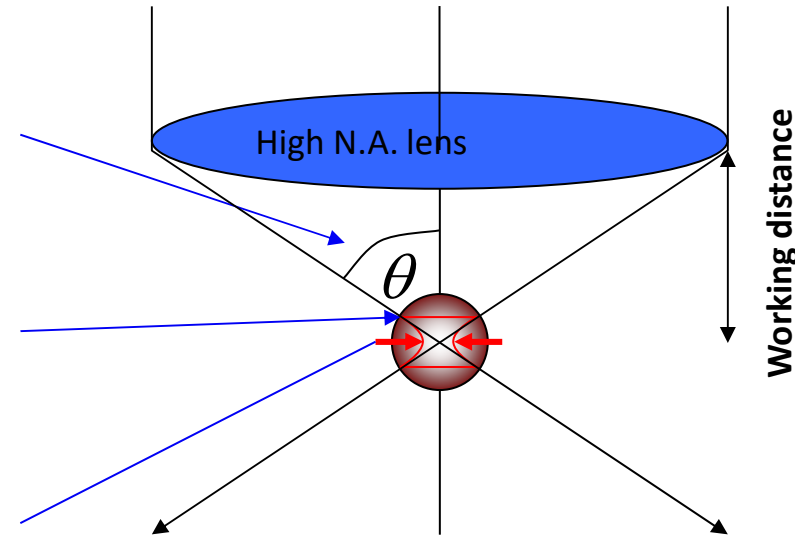
Objective x100



Collection solid angle
 Large for high N.A. lens
 Small for low N.A. lens

Sampling volume
 Small for high N.A. lens
 Large for low N.A. lens

Laser spot size
 Small for high N.A. lens
 Large for low N.A. lens








$$NA = n \cdot \sin(\Theta)$$

n: refractive index

Θ : aperture angle

Collecting lens

Objective

	Objective	N.A.	Working distance [mm]
	x100	0.90	0.21
	x50	0.75	0.38
	x10	0.25	10.6
	x100 LWD	0.80	3.4
	x50 LWD	0.50	10.6

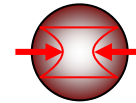
Collecting lens

Objective x100

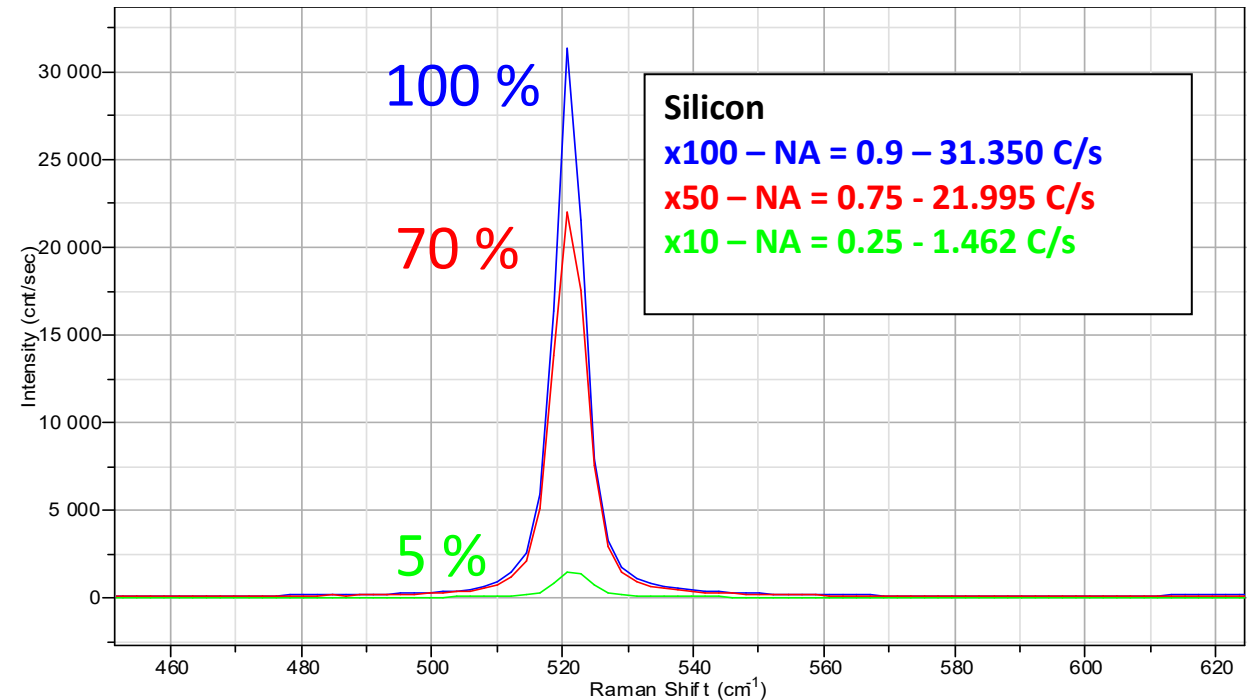
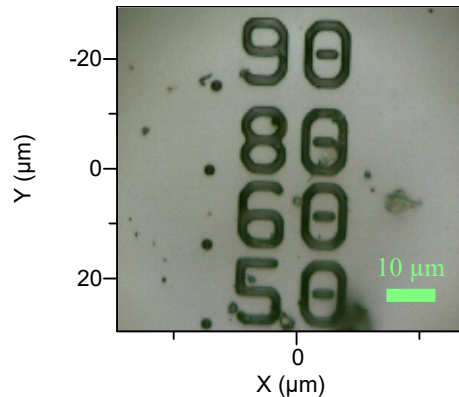
A distinction between opaque and transparent samples must be made

For **opaque samples**, high N.A. lenses work better because there is almost no penetration of the laser into the sample. High N.A. lenses enable:

- High laser power density ($\text{mW}/\mu\text{m}^3$) → **increases sensitivity**
- Wide collection solid angle → **increases sensitivity**



Example for an opaque sample: Silicon wafer



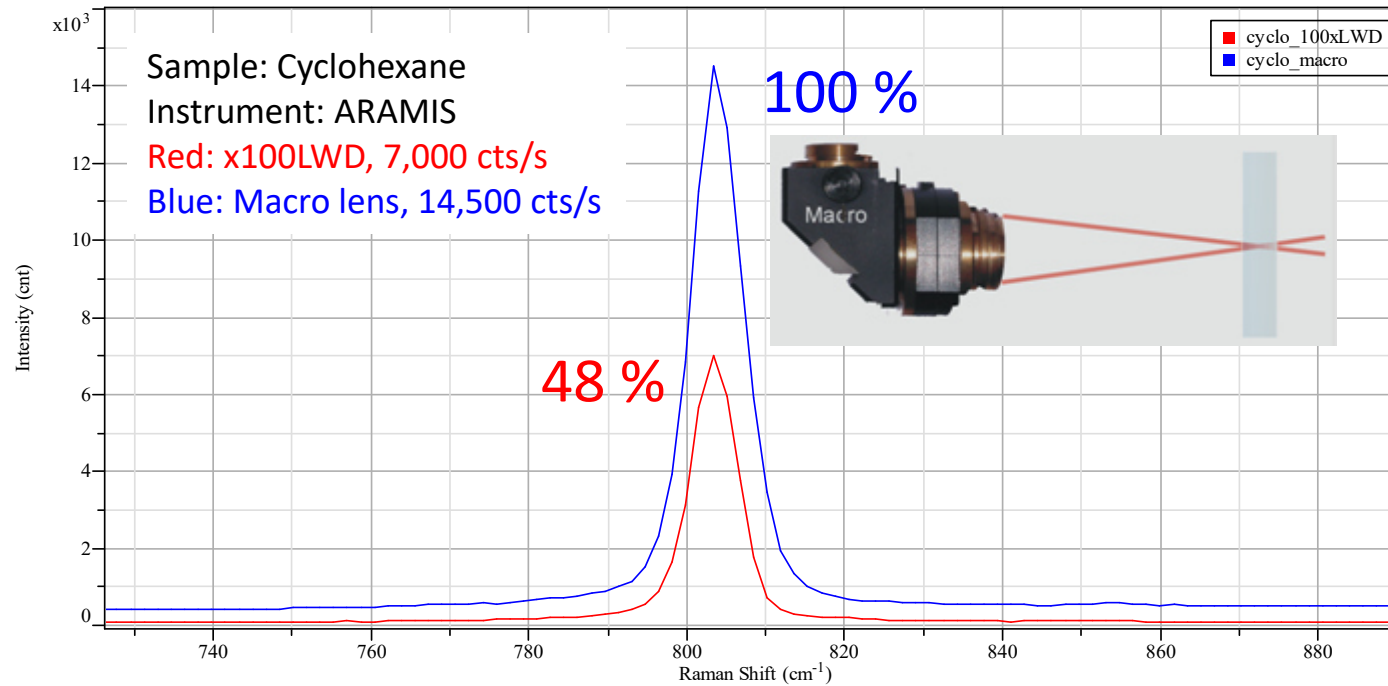
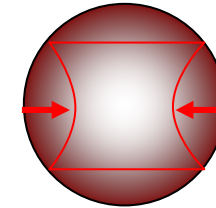
Collecting lens

Objective x100

A distinction between opaque and transparent samples must be made

For transparent samples, low N.A. lenses work better because of increased penetration depth of the laser into the sample. Low N.A. lenses enable:

- Large sampling volume → **increases sensitivity**

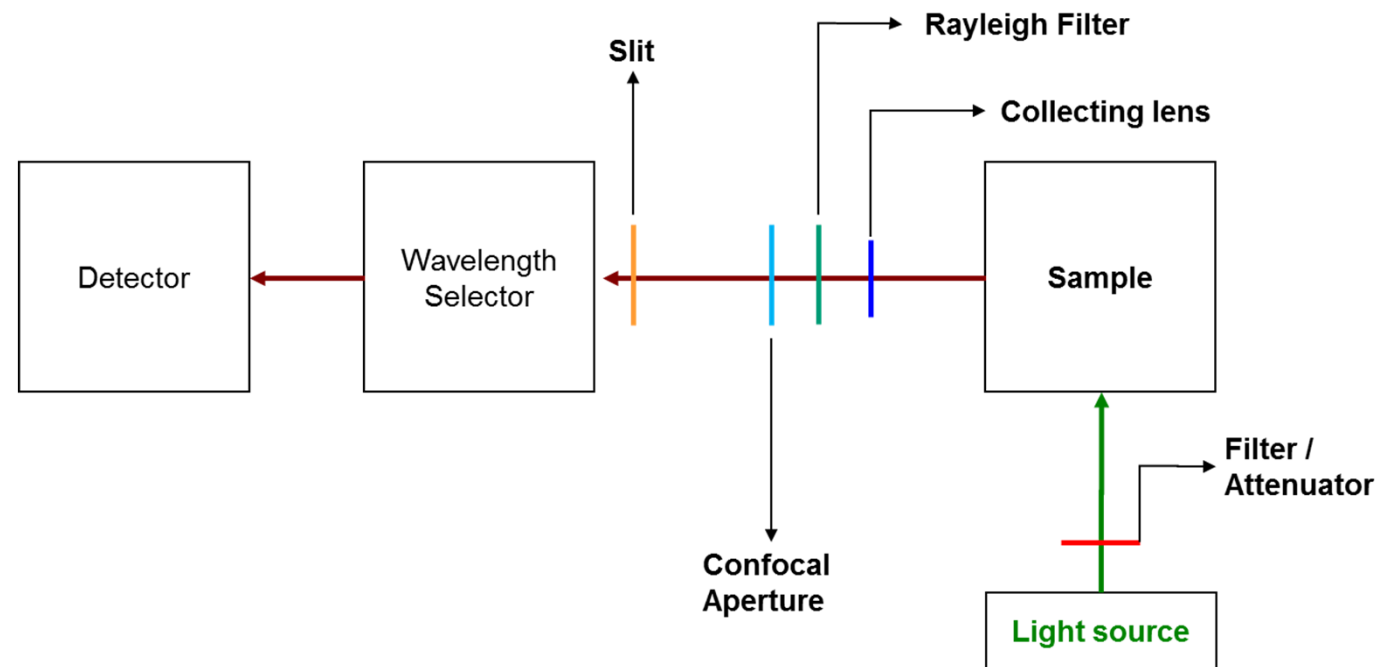


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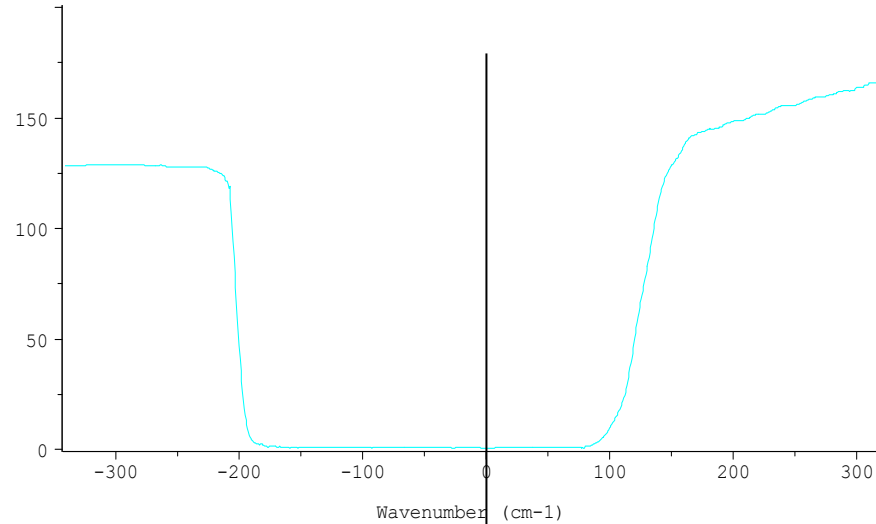
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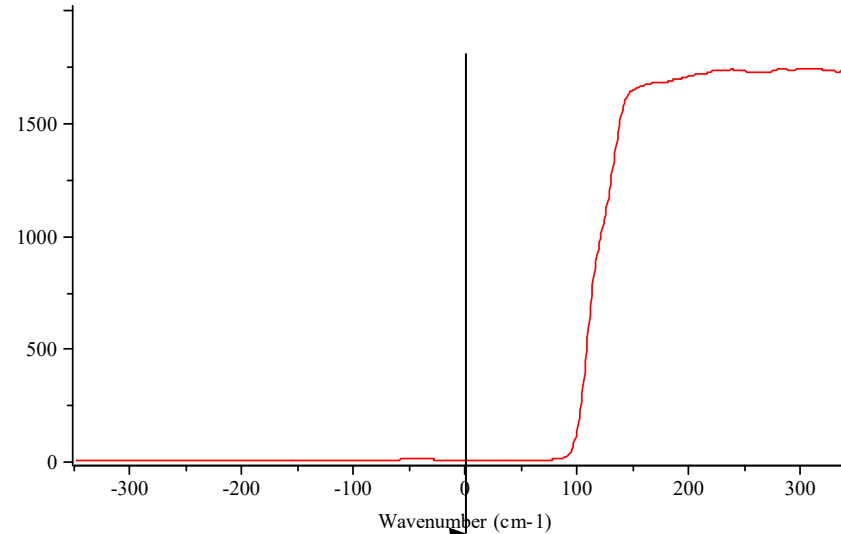


Rayleigh Filter

White light spectrum with a notch filter



White light spectrum with an edge filter

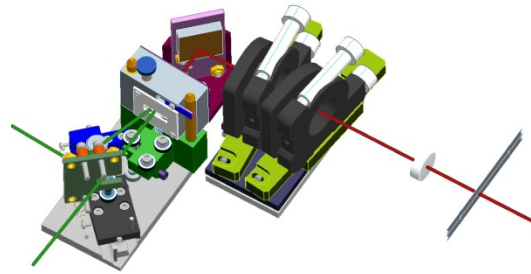


Zero Raman shift → Excitation laser position

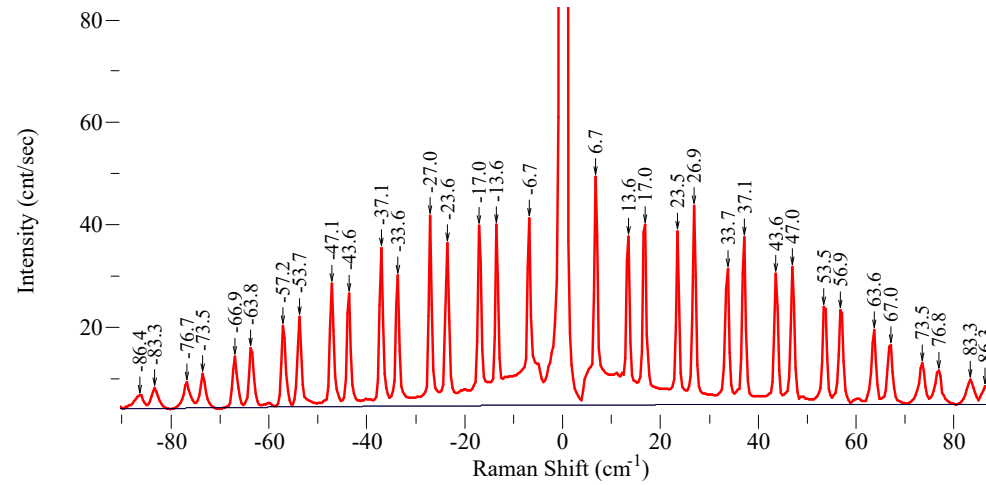
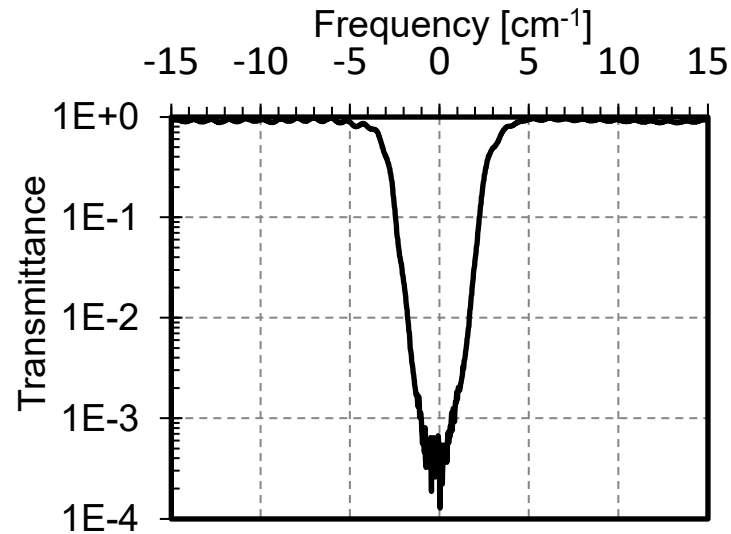
- | | |
|--|---|
| <ul style="list-style-type: none"> ■ A finite life time ■ Stokes and Anti-Stokes Raman | <ul style="list-style-type: none"> ■ A virtually infinite life time ■ Stokes Raman only |
|--|---|

There is a cost advantages to the edge filter due to **no aging**, but Anti-Stokes-Raman is **not** obtainable

Rayleigh Filter



ULF (Ultra Low Frequency)



Ultra-Lattice of GeSi

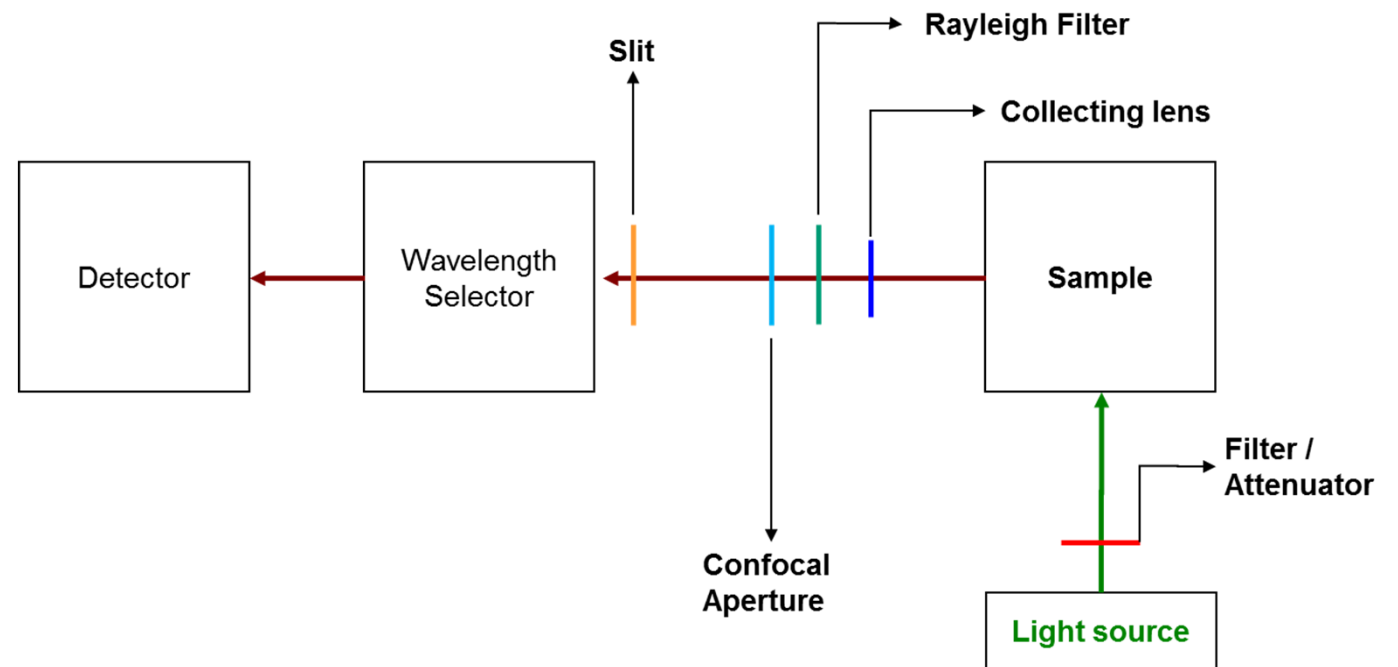
data courtesy of :
P. H. Tan, State Key Laboratory for SL and Microstr., Institute of Semiconductors, Beijing, P. R. China
K. Brunner, University Wurzburg, Germany

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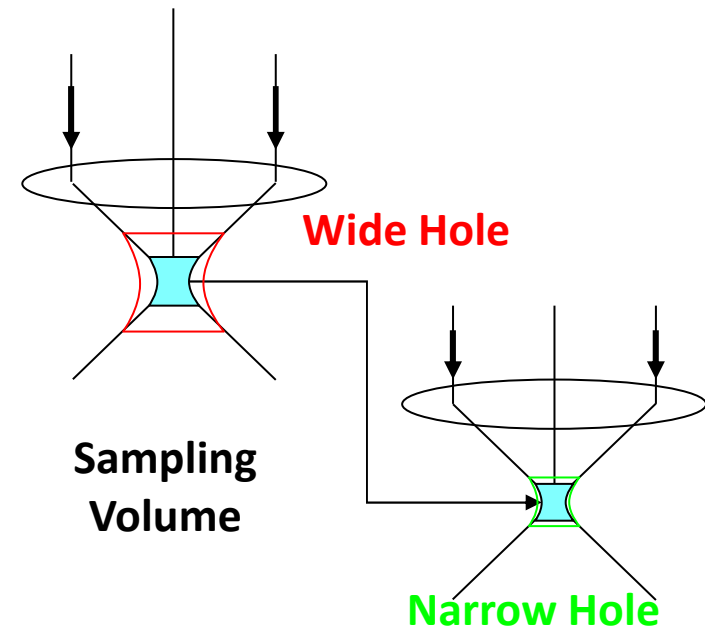
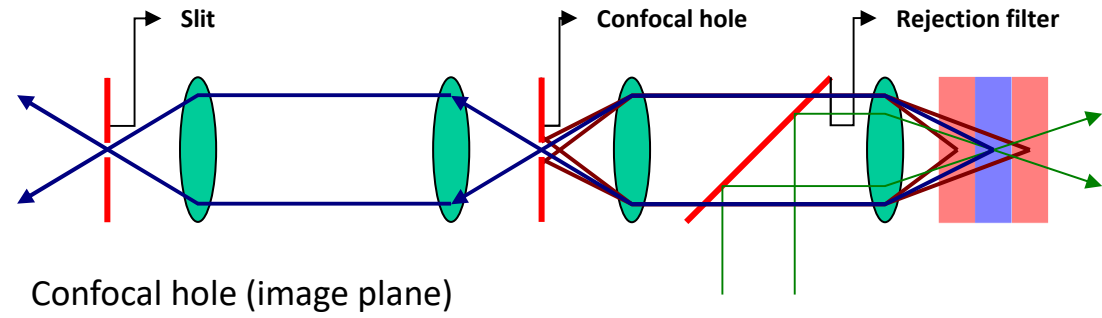
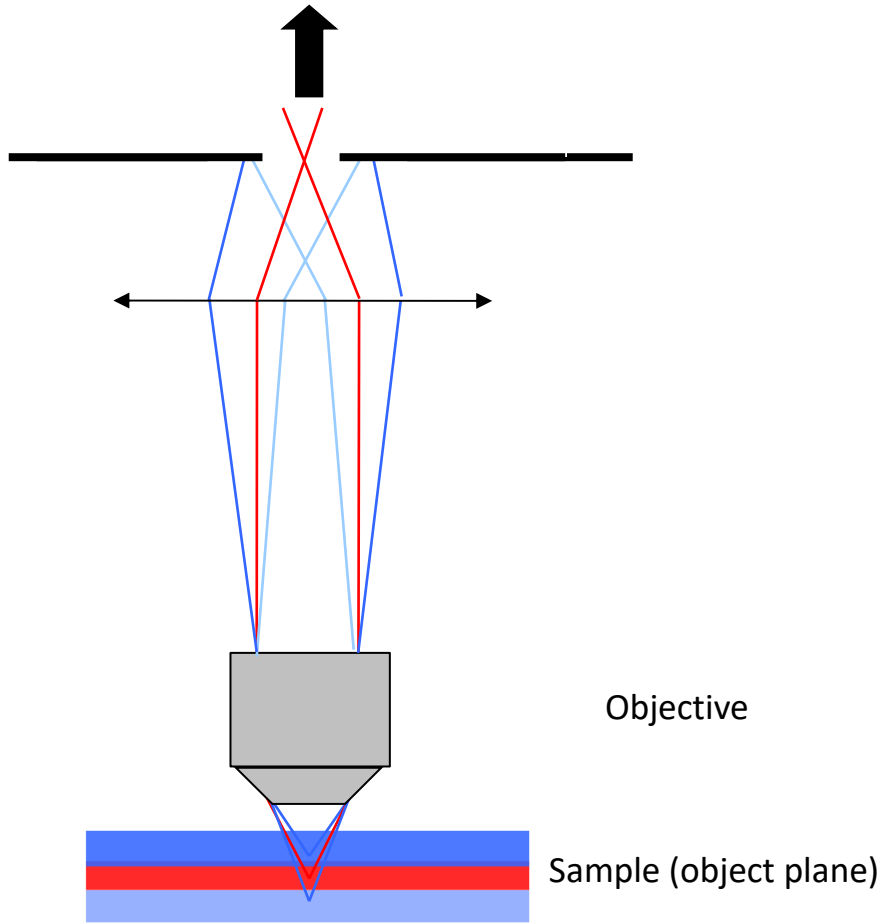
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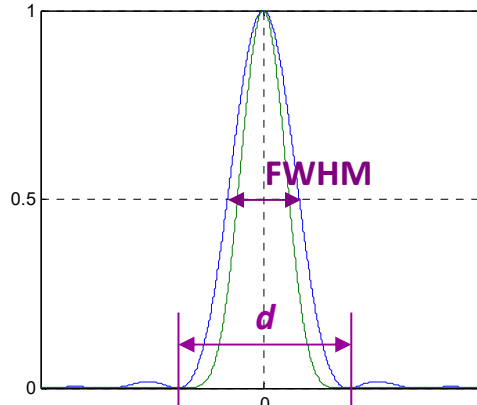
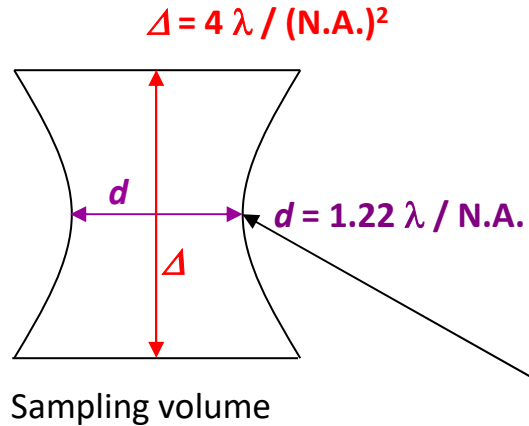
Confocal Aperture

Hole 1000



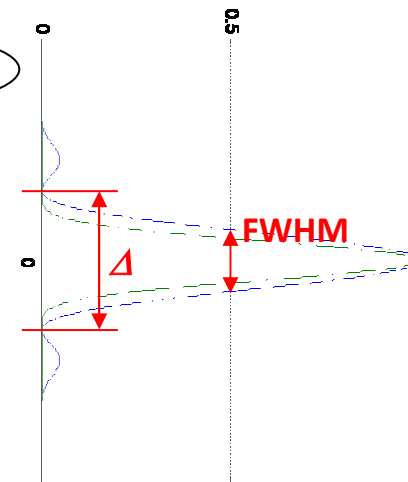
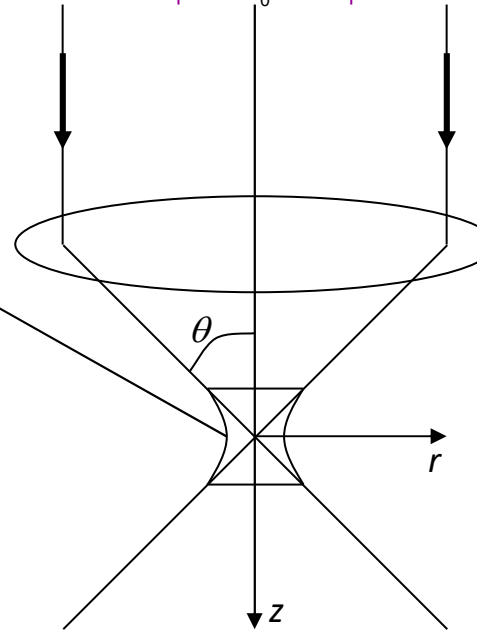
Confocal Aperture

Conventional vs. Confocal



Lateral Resolution
 $D_{xy} = 0.59 \lambda M^2 / NA$

Typical spatial resolution, if the
 confocal hole aperture fits with the
 spot diameter



$$D_z = \frac{0.88 \lambda M^2}{n - \sqrt{n^2 - NA^2}}$$

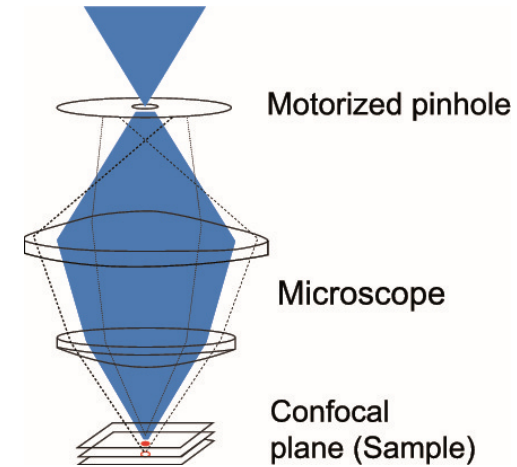
Confocal Aperture

Hole

Theoretical limits in spatial resolution

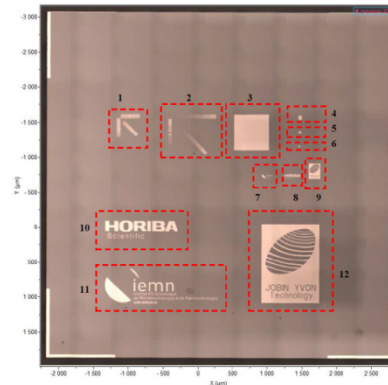
	D_{xy} (μm)	D_z (μm)
473	0,3	0,9
532	0,4	1,0
633	0,5	1,2
785	0,6	1,5

With 100x/0,9 objective, in air, and $M^2=1.1$

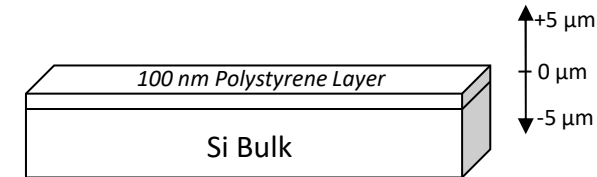


Experimental determined limits in spatial resolution done with 532 nm laser & x100 dry air objective

For lateral resolution:
sample from IEMN laboratory
40 nm gold on Si 1:1:1



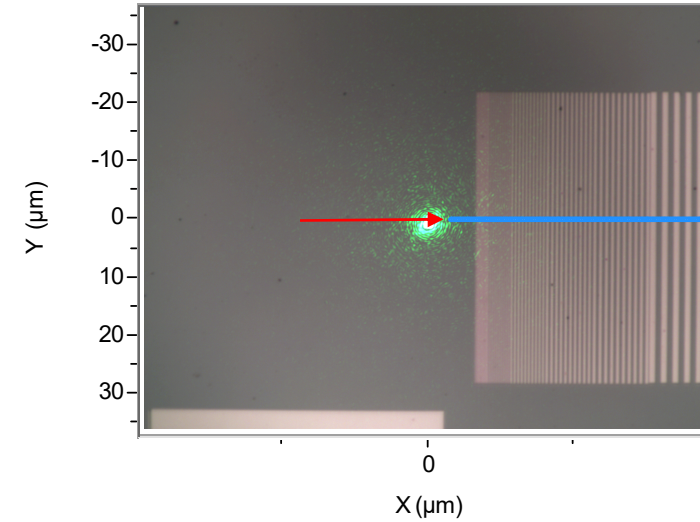
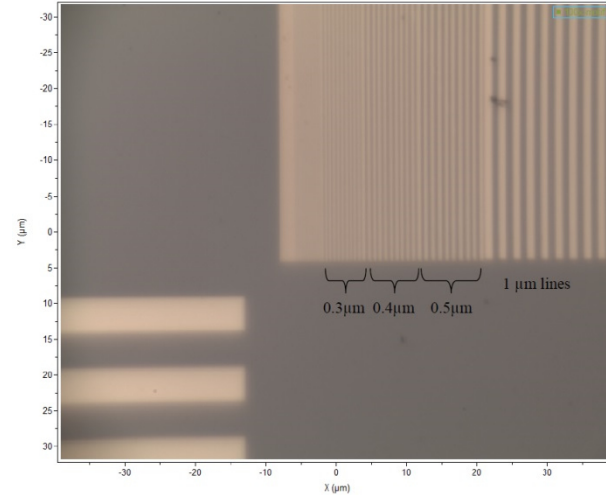
For axial resolution:
100 nm PS layer on Si
1:1:1



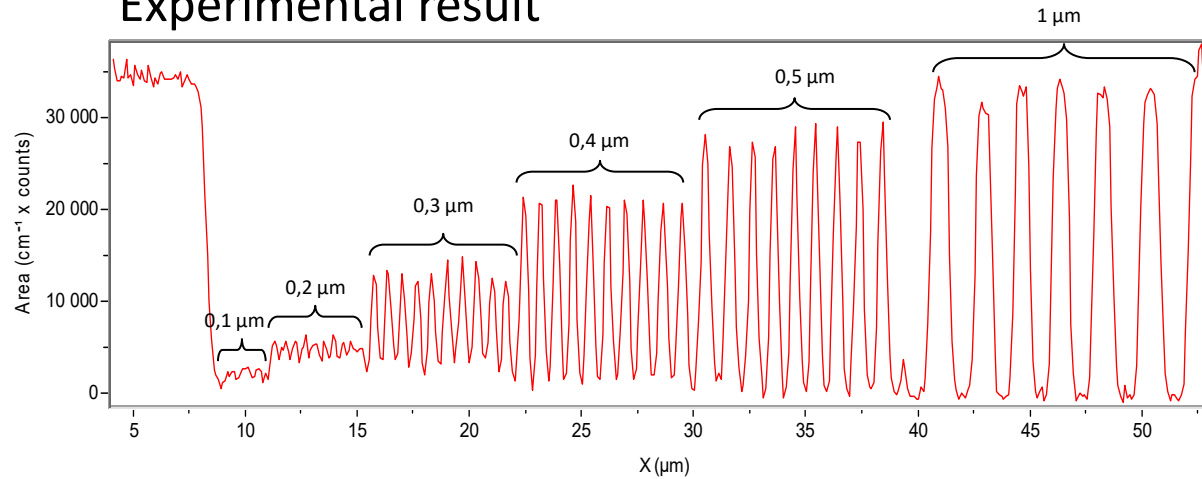
Confocal Aperture

Hole 1000

Lateral resolution:
Line scan over gold
pattern (40 nm
height) with varying
distance



Experimental result

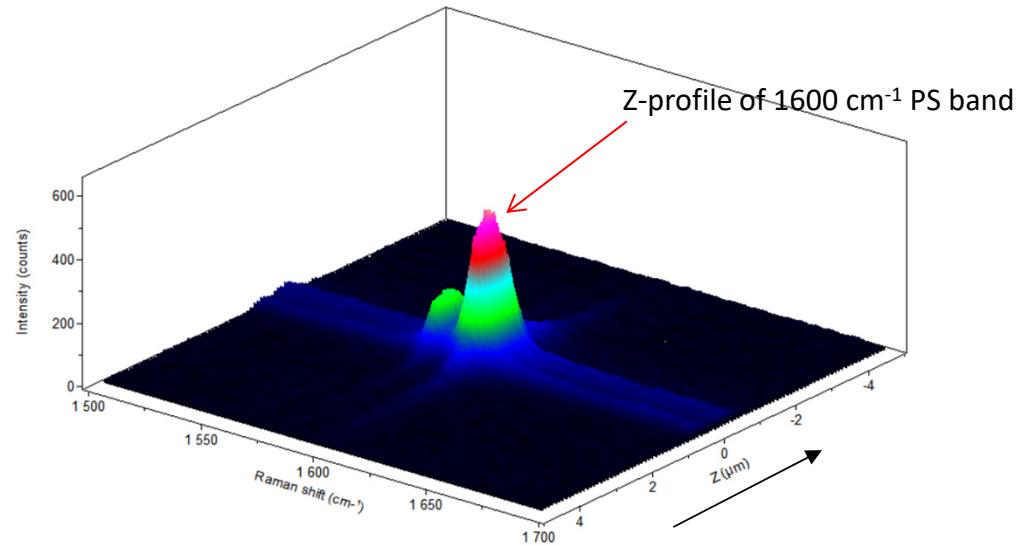
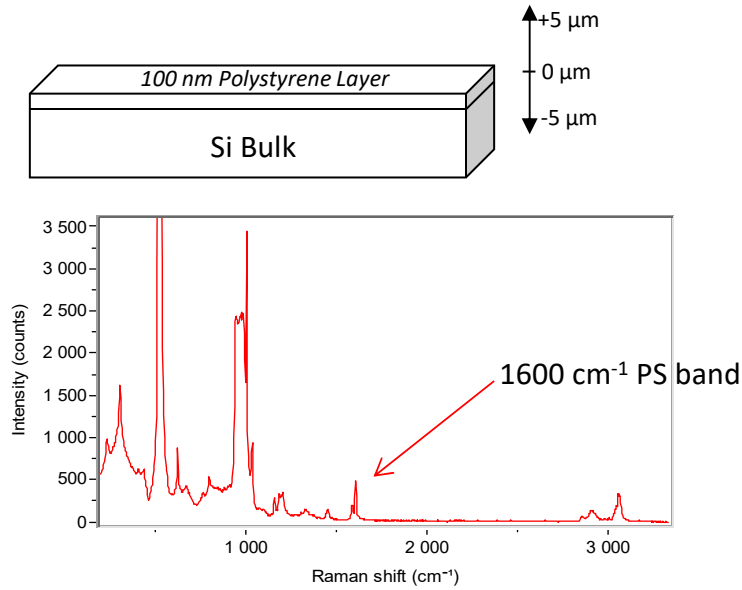


Theory	D_{xy} (μm)	D_z (μm)
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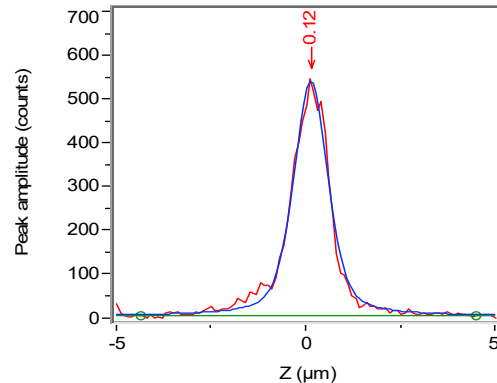
Confocal Aperture

Hole 1000

Axial resolution: z-scan through a 100 nm thick Polystyrene layer



Experimental result: Fit of z-profile of 1600 cm⁻¹ PS band



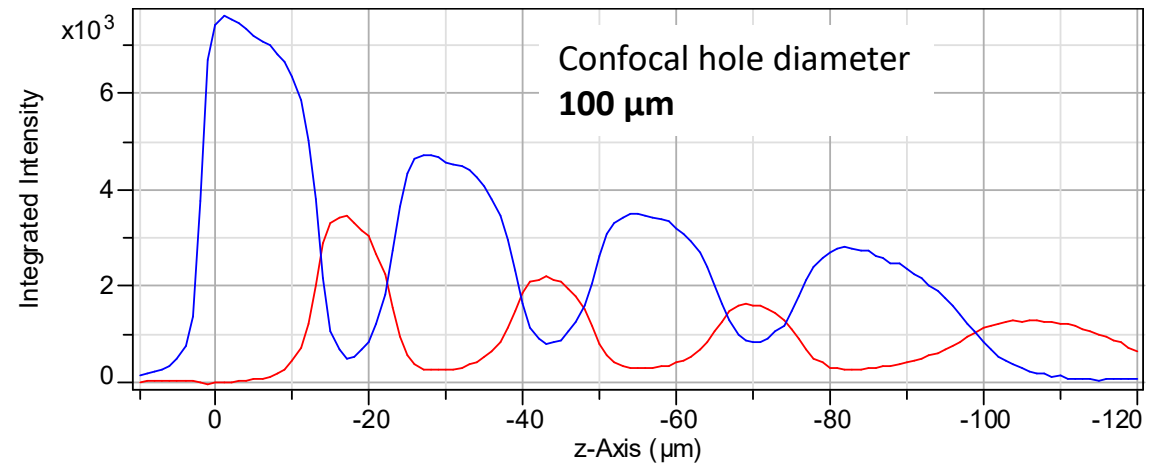
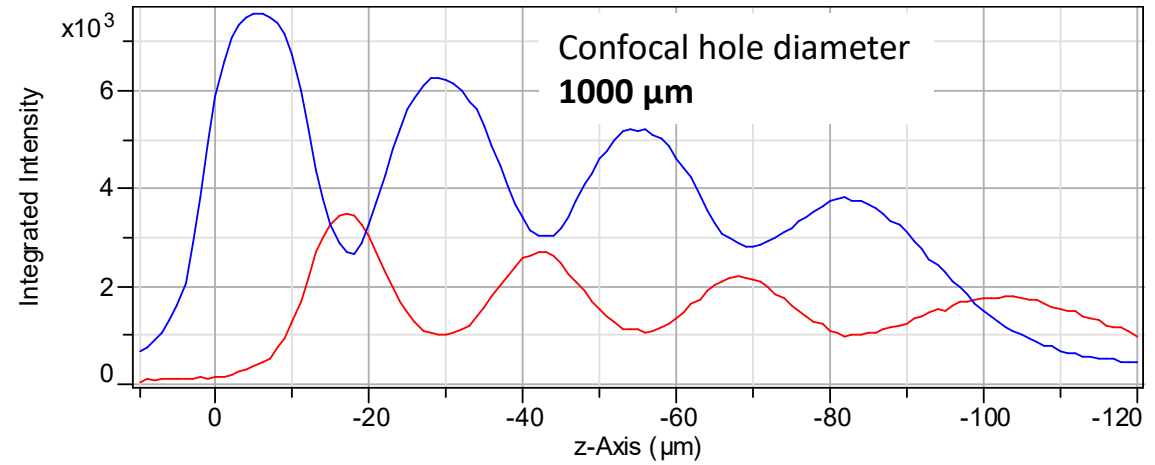
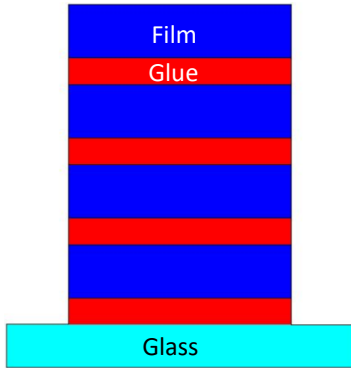
Peakposition	Amplitude	Gauß-Lorenz	FWHM	Area
0,12	537,62	0,56	1,05	701,58

Theory	D _{xy} (μm)	D _z (μm)
473	0,3	0,9
532	0,4	1,0
633	0,5	1,2
785	0,6	1,5

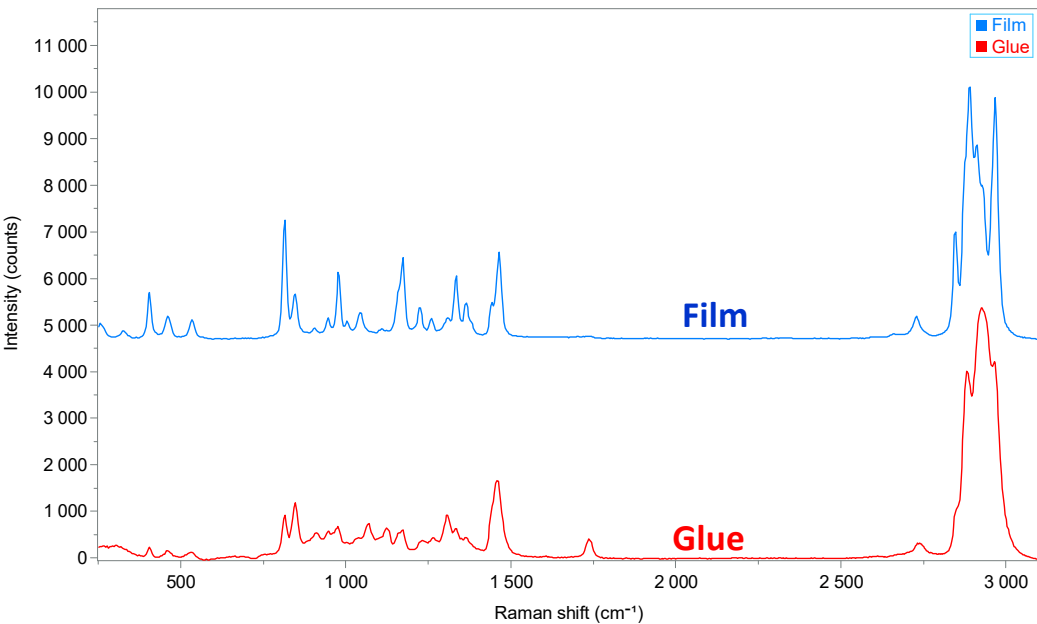
Confocal Aperture

Hole 1000

Z-scan of multi layered polymer



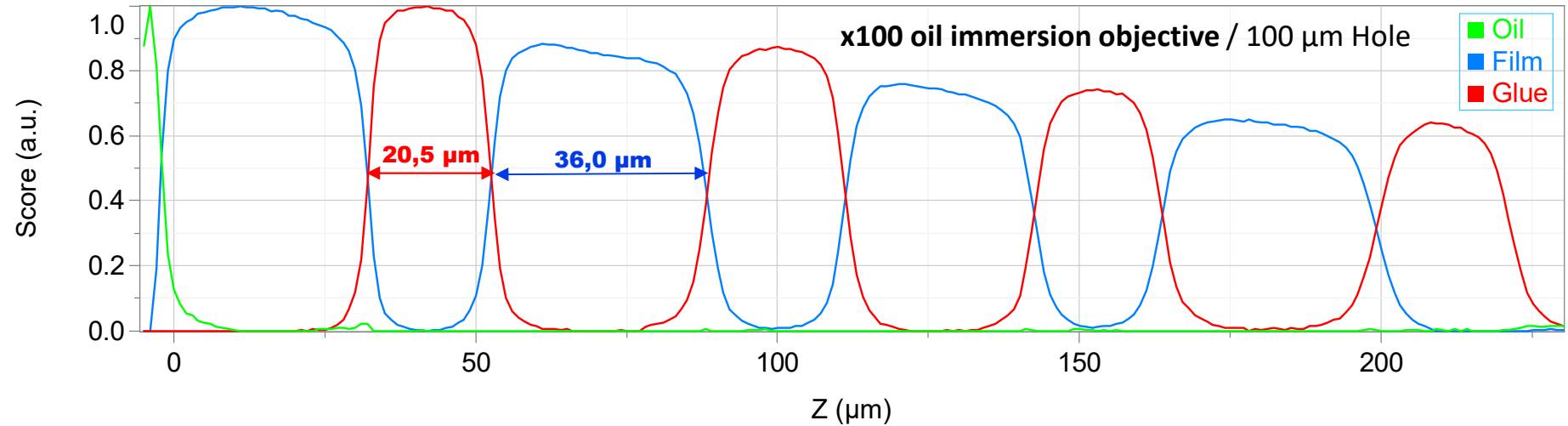
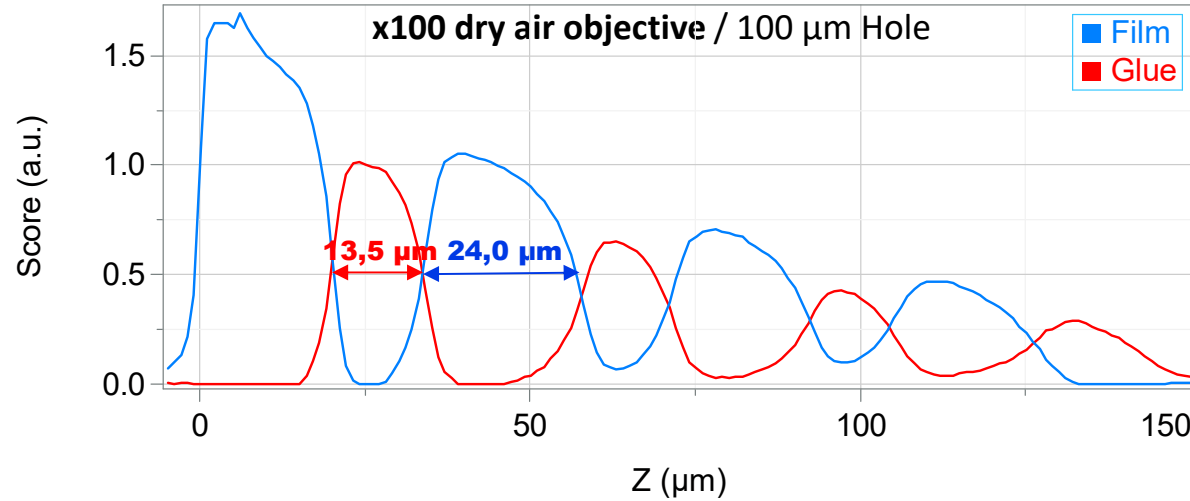
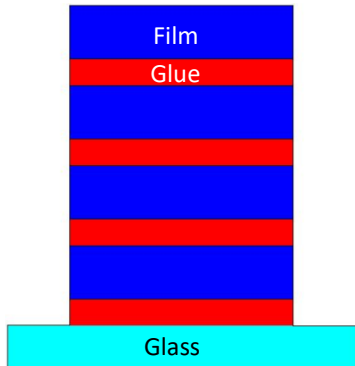
The narrow hole improves axial resolution.



Confocal Aperture

Confocal z-scan x100 dry air objective versus x100 oil immersion objective

The oil immersion objective improves axial resolution remarkably in comparison with the x100 dry air objective



Confocal Aperture

In absence of refraction ($n_1 = n_2$) all rays are focused at point P_0 , which is related to the nominal focal distance z_0 (= corresponding to the value on the micrometer scale of microscope).

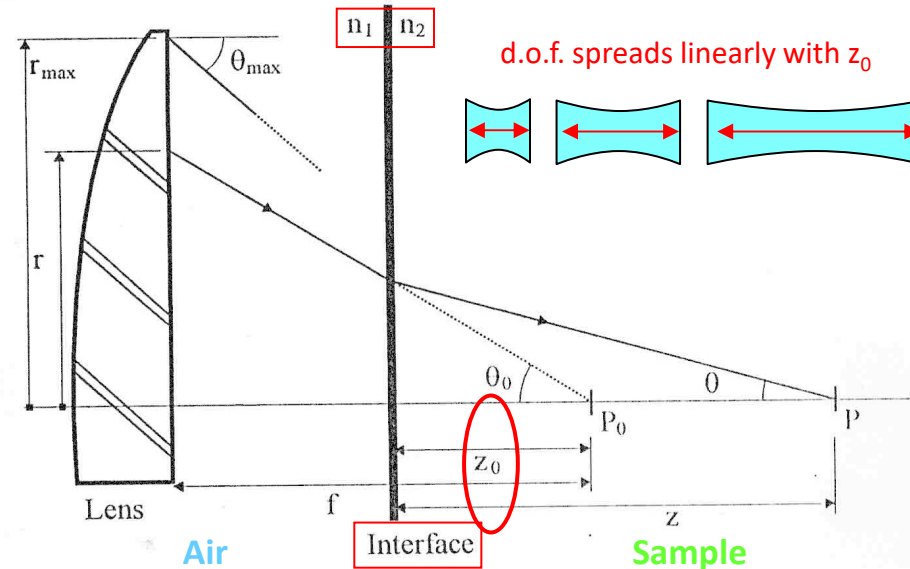
But refraction ($n_1 \neq n_2$) causes:

Rays to be focused deeper into the sample at point P which lies a distance z below the sample/air interface, the actual point of focus z doesn't correspond with the nominal focal point z_0 anymore.

The length of focus (or depth of focus = d.o.f) increases linearly with z_0 , means the length of laser focus spreads with increasing depth: $d.o.f = k \cdot z_0$



Data and figures from
Modeling and Measuring the Effect of Refraction on the Depth Resolution of
Confocal Raman microscopy
Neil. J. Overall, Applied Spectroscopy, Vol. 54, 6, 2000

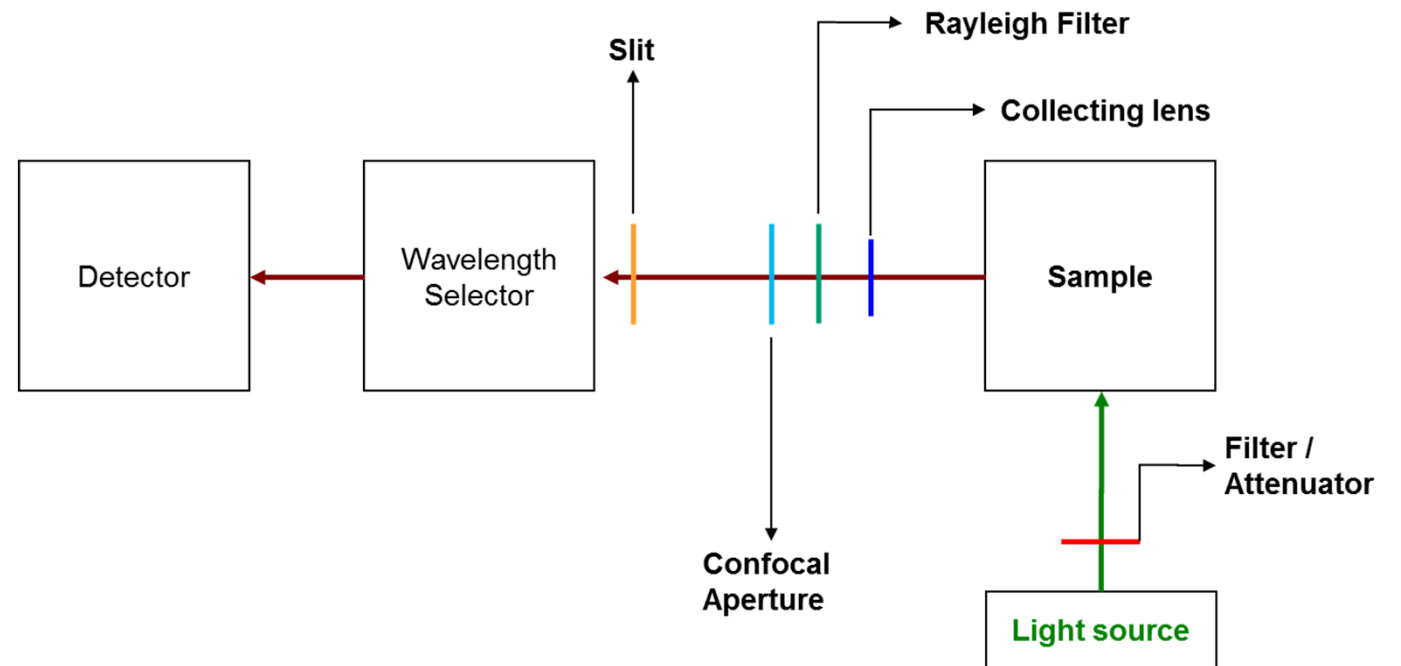


Outline

- Principles of Raman effect in a nutshell
- What information does Raman give?

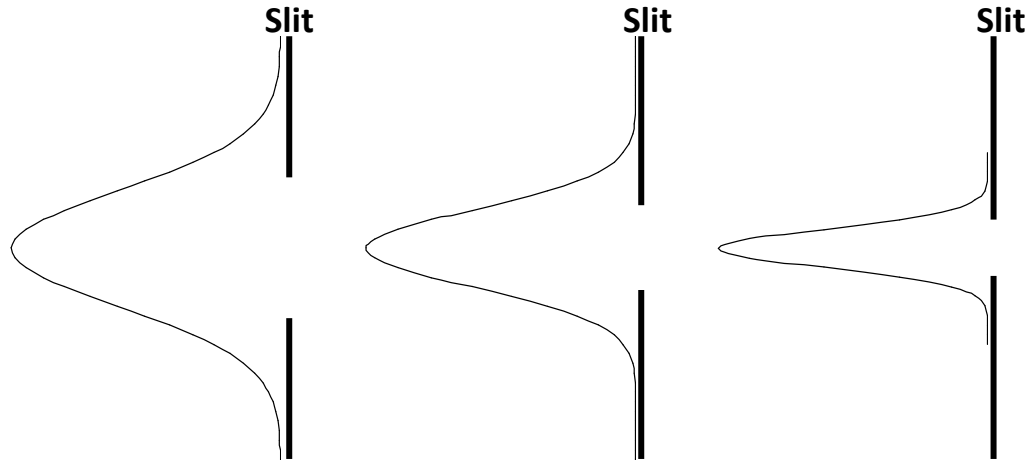
■ Instrument Parameters I

- Laser
- Filter / Attenuator
- Collecting lens
- Rayleigh Filter
- Confocal Aperture
- Slit
- Spectrometer & Grating
- Detector & Quantum Efficiency
- Correction of Instrument Function



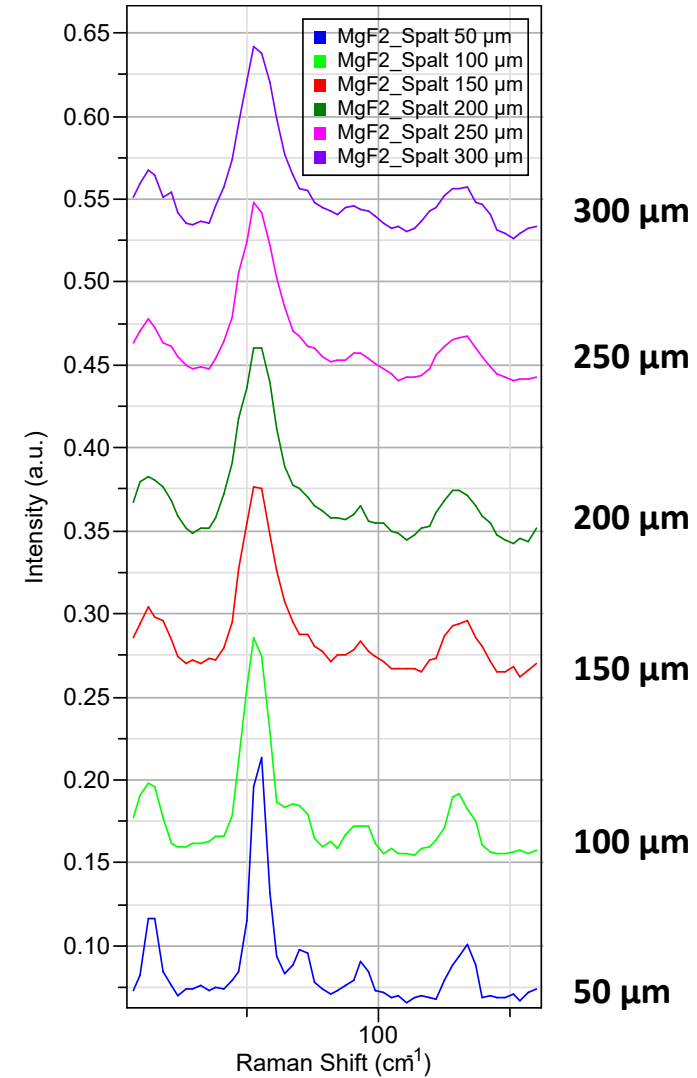
Slit

Slit 100



One of the parameters that determines the spectral resolution is the entrance slit width. The narrower the slit, the narrower the FWHM (full width at half maximum), and the higher the spectral resolution.

When recording a line with a natural width smaller than the monochromator's resolution, the measured width will reflect the spectrograph's resolution.

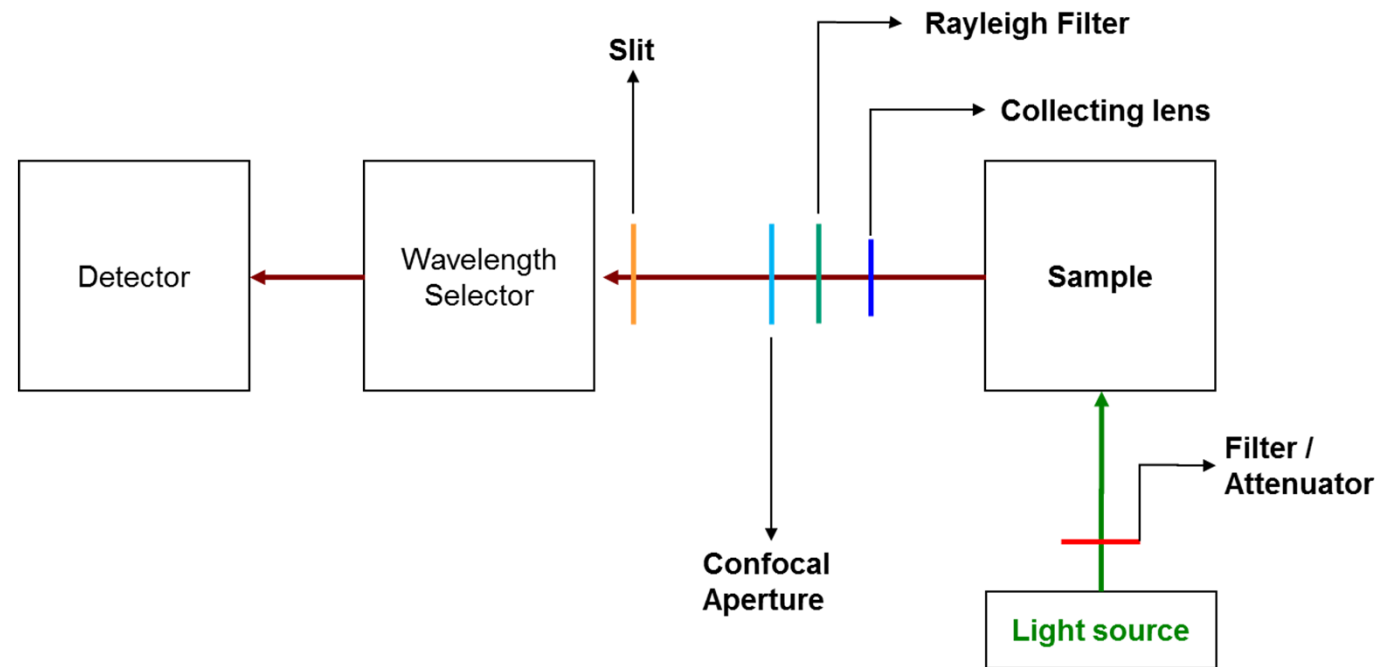


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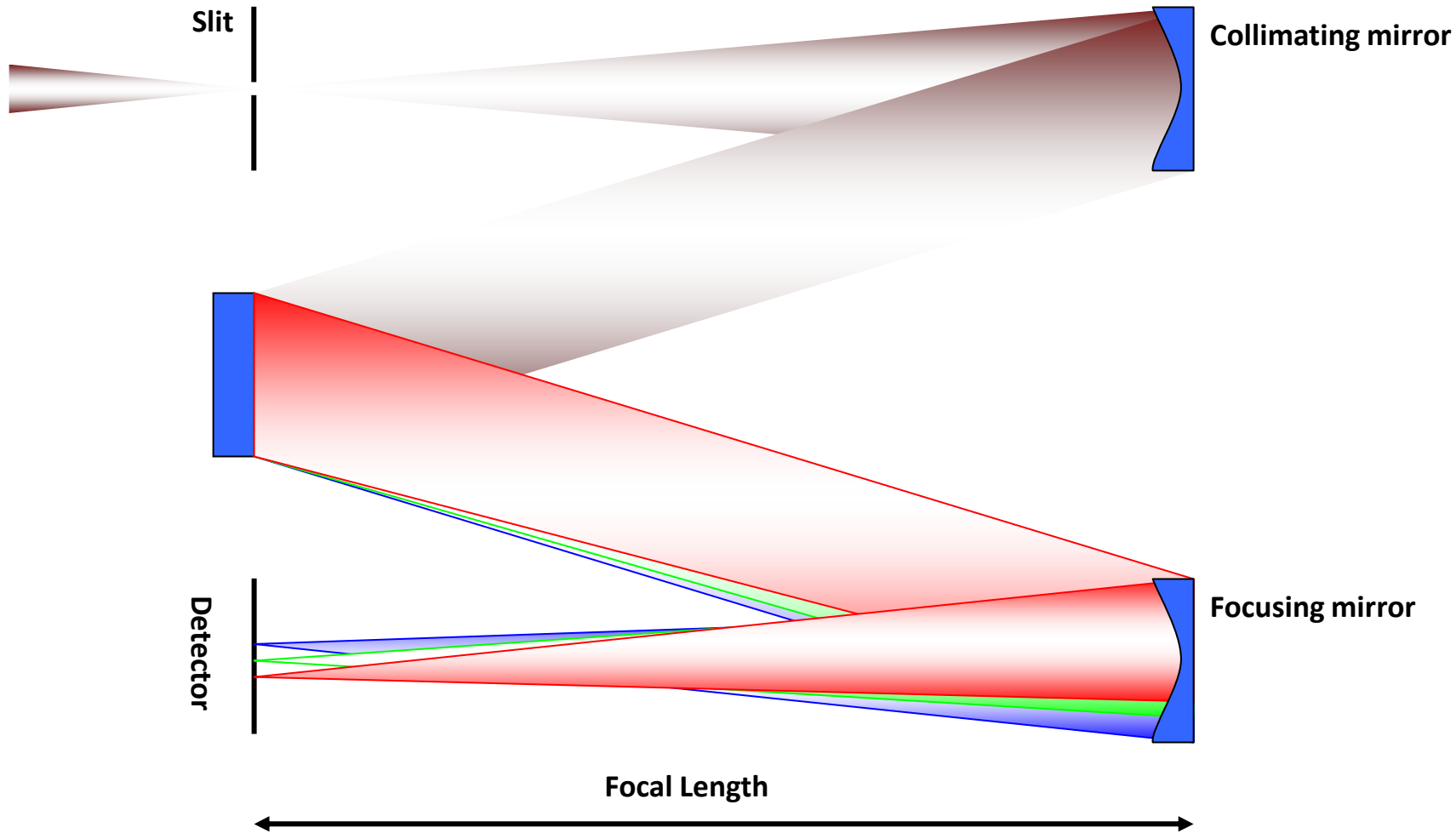
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Spectrometer & Grating

Spectro (cm^{-1})	520.05	
Range	150	3400
Grating	1800 gr/mm	

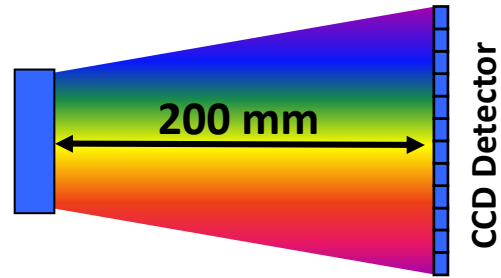


Schematic diagram of a Czerny-Turner spectrograph

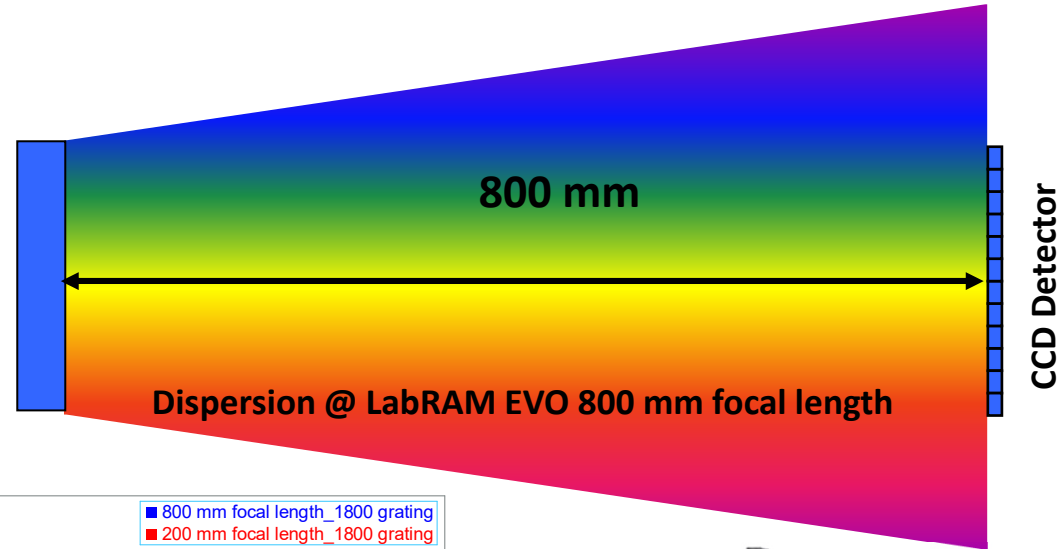
Spectrometer & Grating

Spectro (cm^{-1})	520.05	
Range	150	3400
Grating	1800 gr/mm	

- Different focal length
- Same grating
- Same excitation wavelength



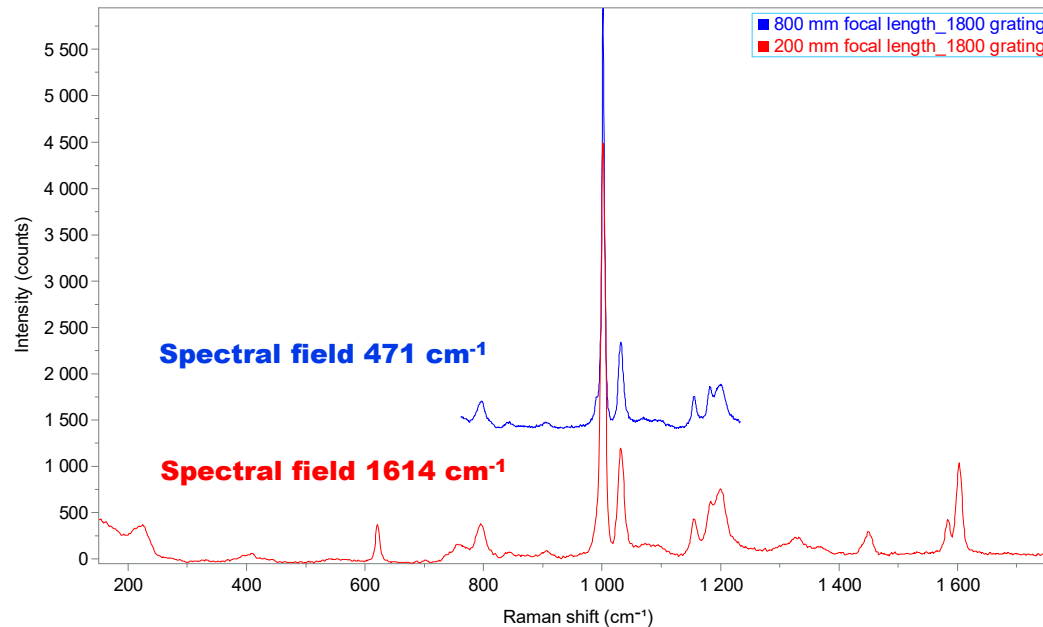
Dispersion @ XploRA 200 mm focal length



Dispersion @ LabRAM EVO 800 mm focal length



200 mm focal length

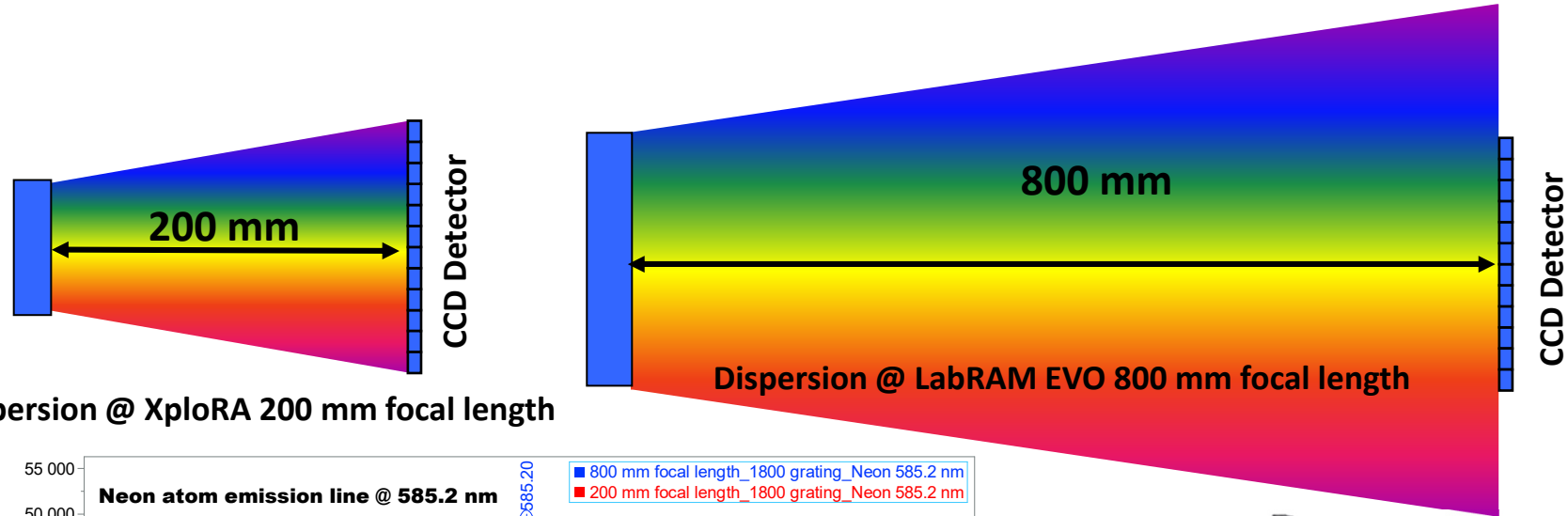


800 mm focal length

Spectrometer & Grating

Spectro (cm^{-1})	520.05	
Range	150	3400
Grating	1800 gr/mm	

- Different focal length
- Same grating
- Same excitation wavelength

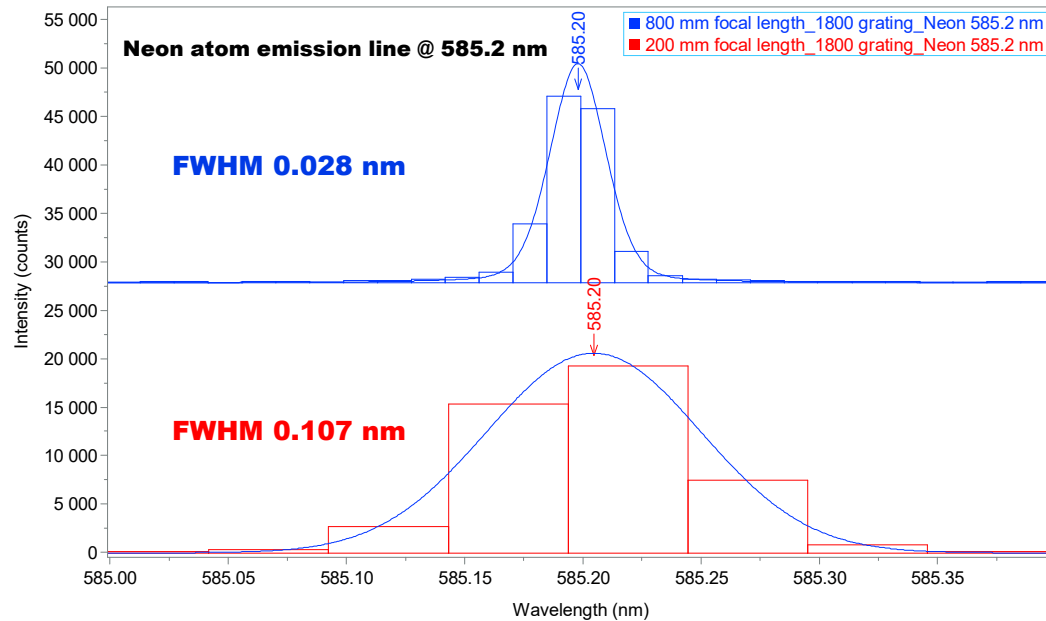


Dispersion @ XploRA 200 mm focal length

Dispersion @ LabRAM EVO 800 mm focal length



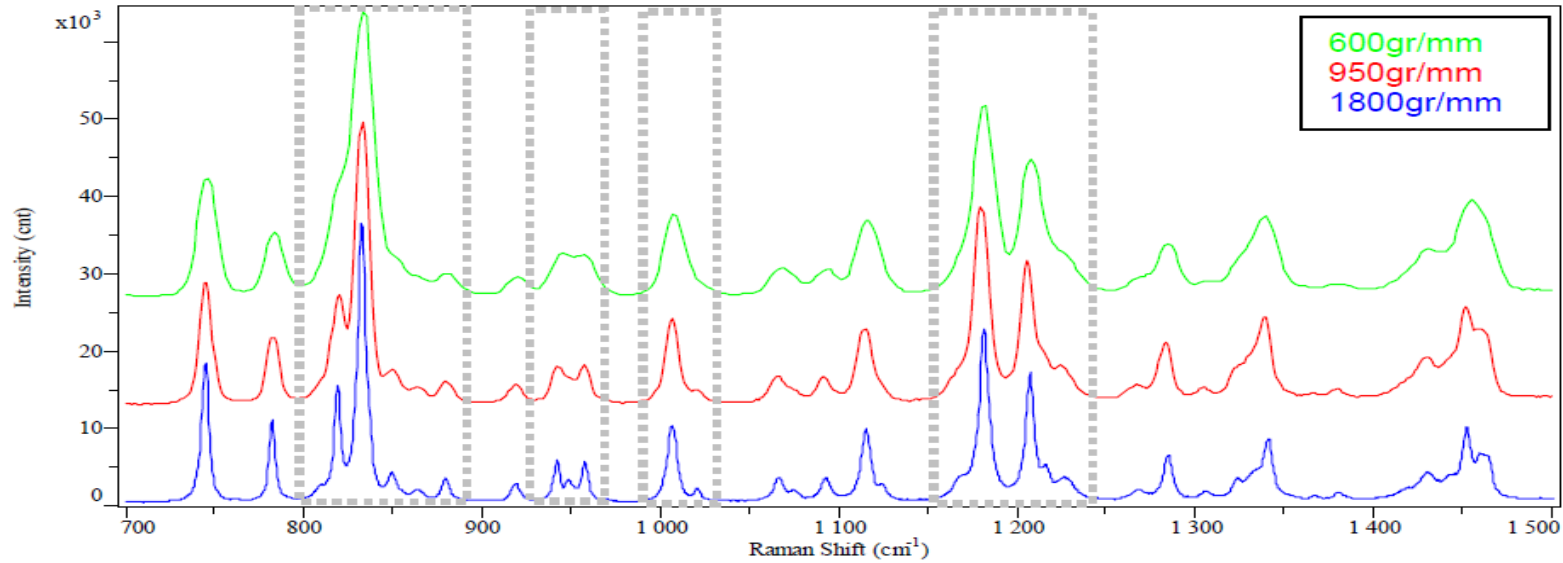
200 mm focal length



800 mm focal length

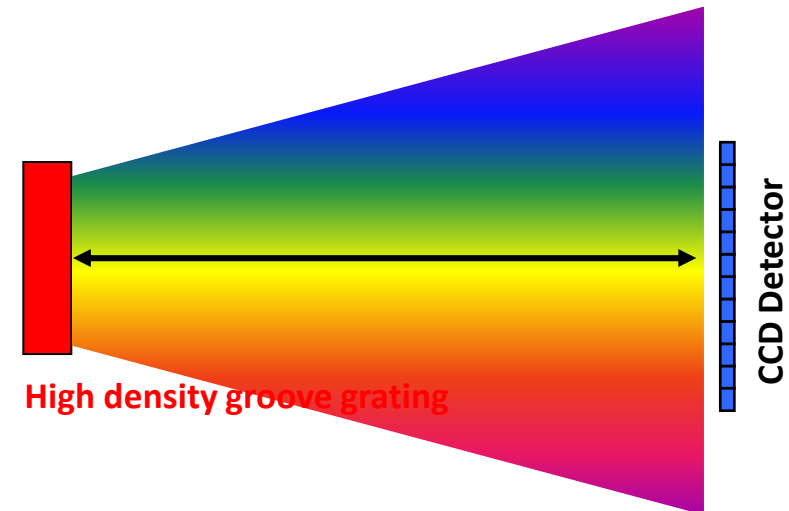
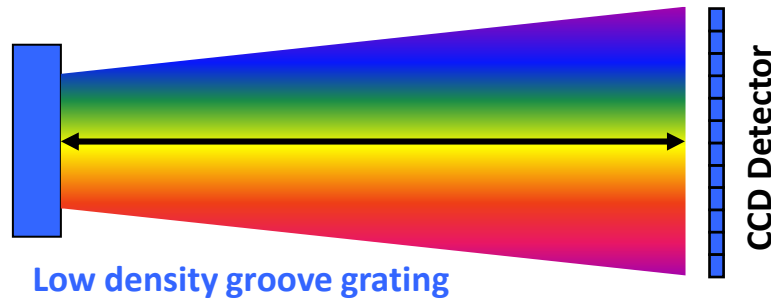
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Spectro (cm^{-1})	520.05	
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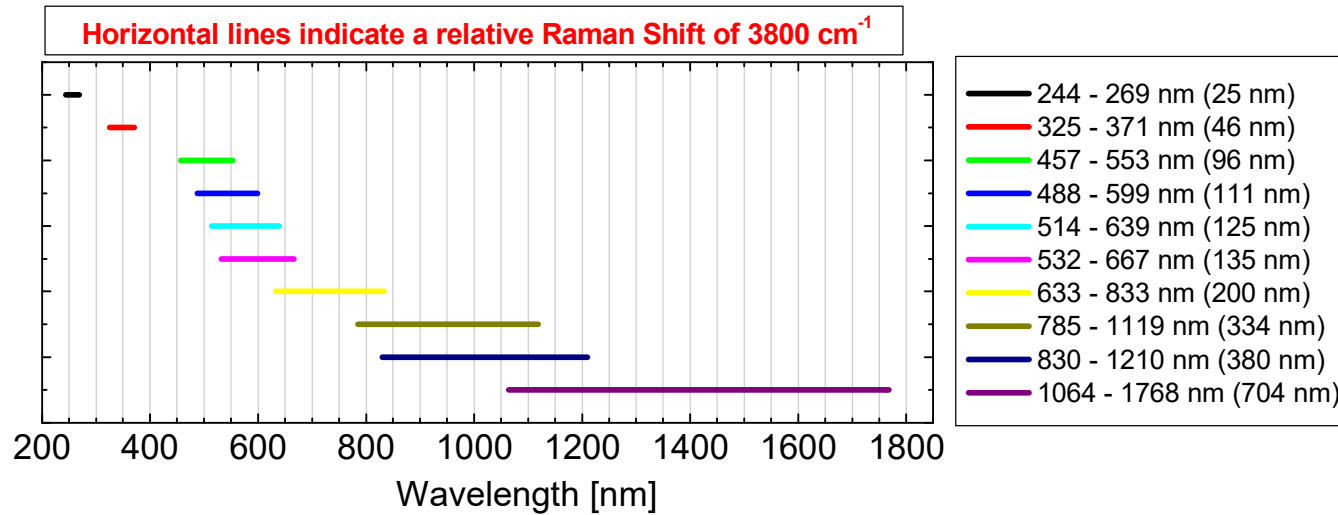
Dispersion as a function of groove density

- **Different grating**
- **Same focal length**
- **Same excitation wavelength**



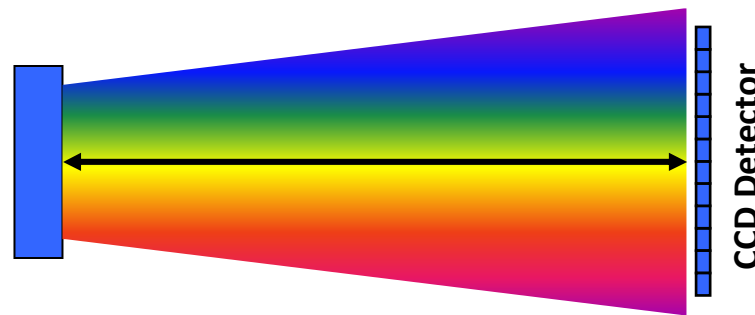
Spectrometer & Grating

Spectro (cm^{-1})	520.05	
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Grating	1800 gr/mm	

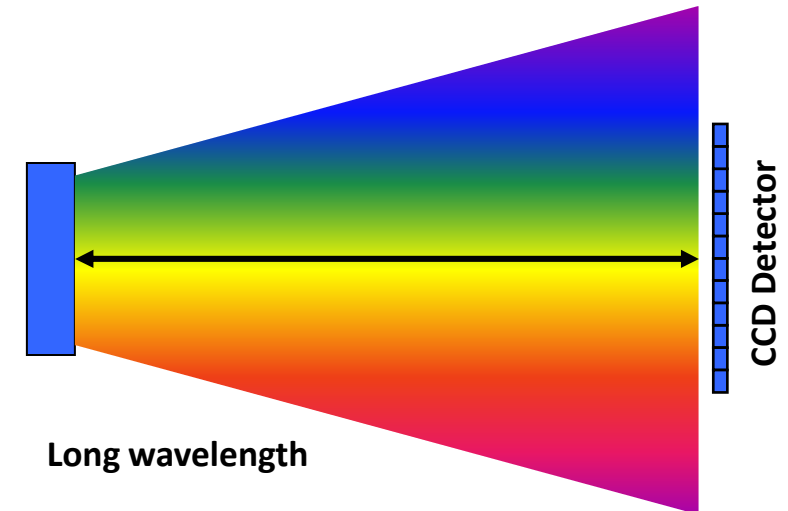


Dispersion as a function of excitation wavelength

- Different excitation wavelength
- Same grating
- Same focal length



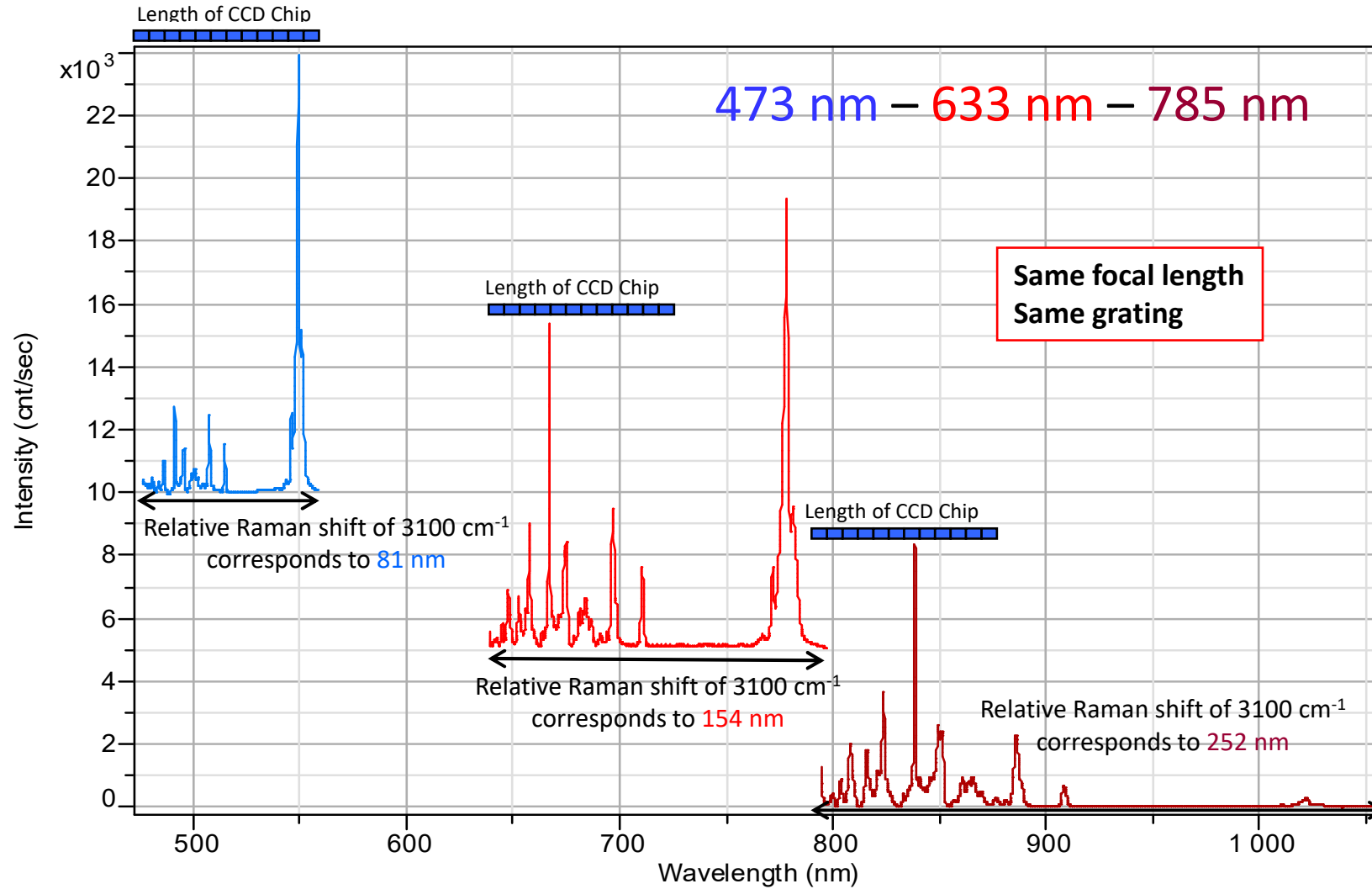
Short wavelength



Long wavelength

Spectrometer & Grating

Spectro (cm^{-1})	520.05	
Range	150	3400
Grating	1800 gr/mm	

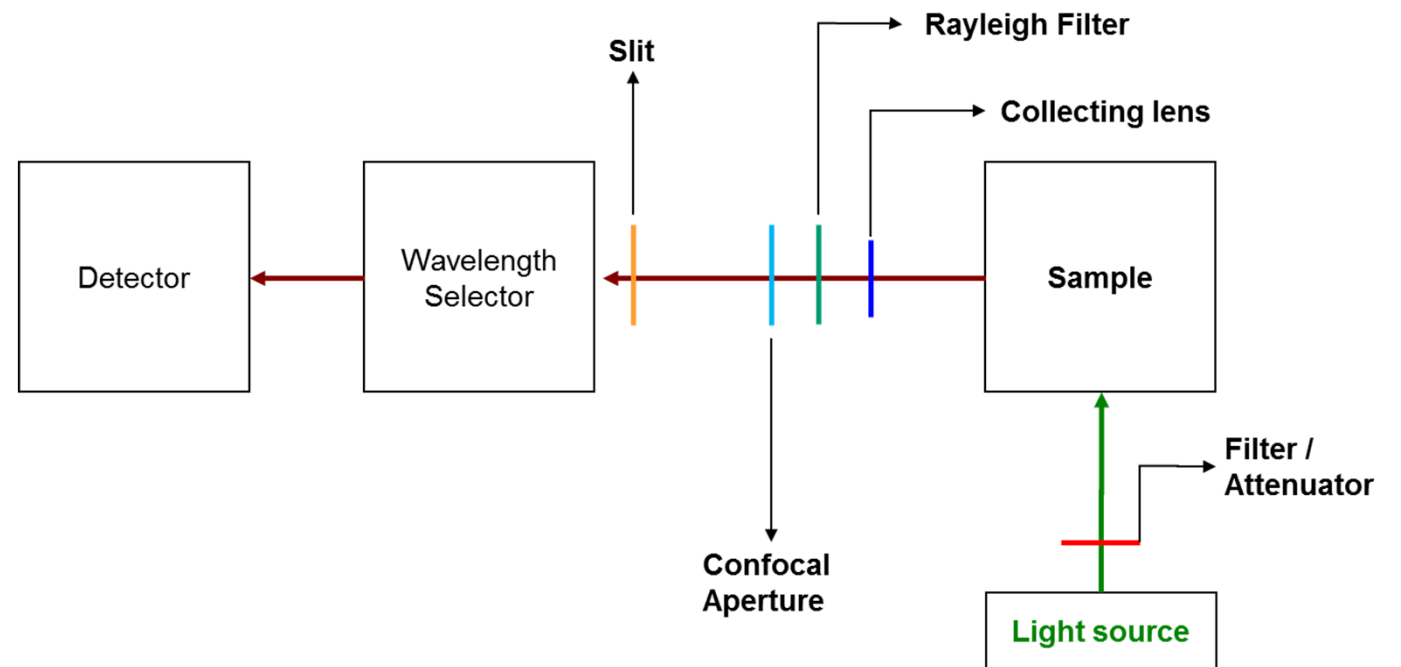


Outline

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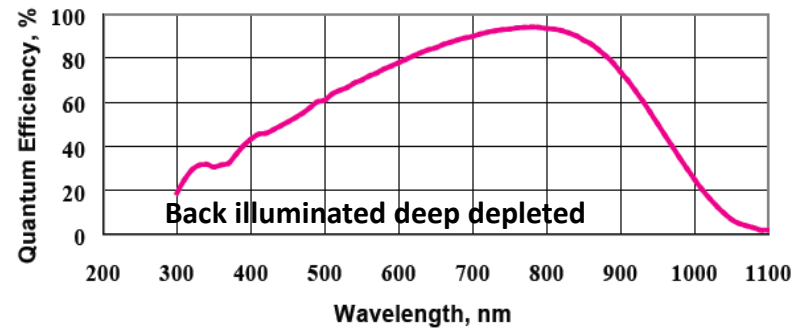
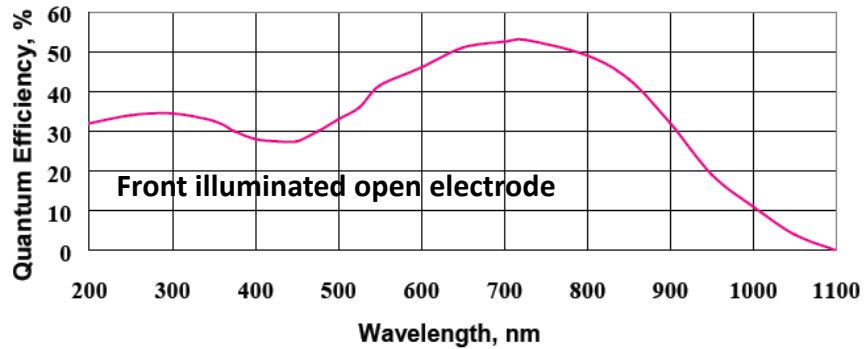
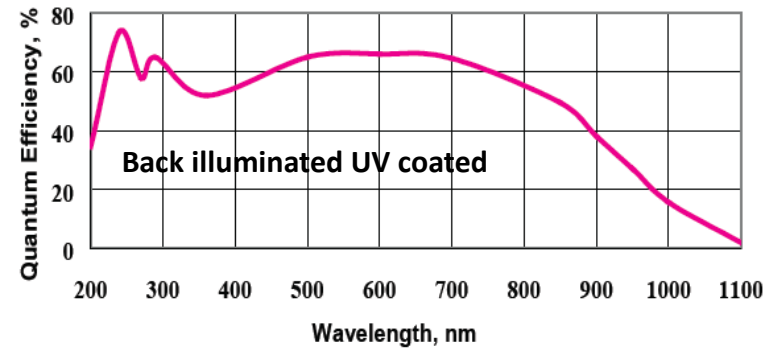
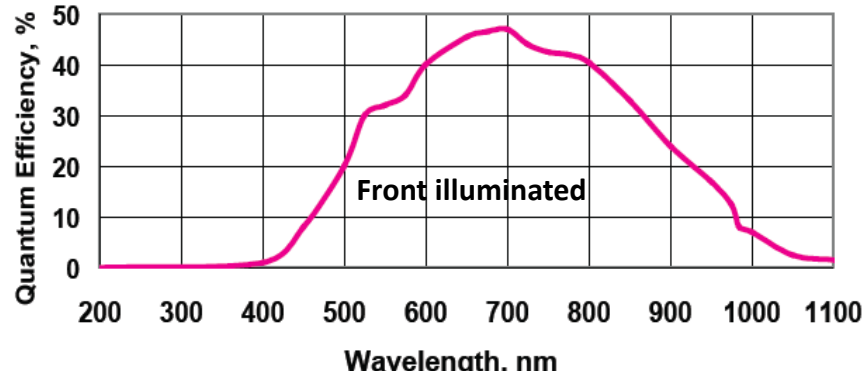
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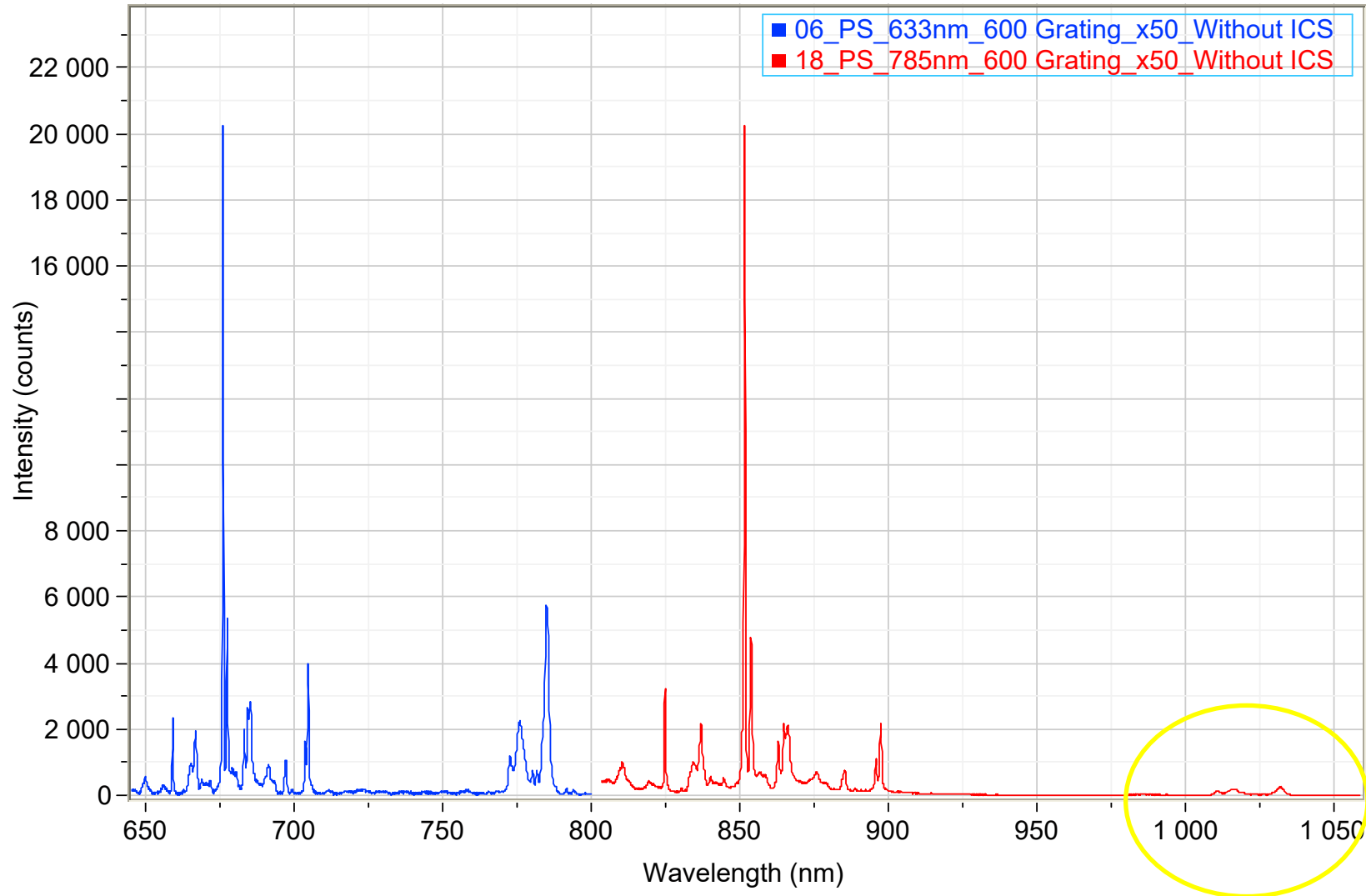


Detector & Quantum Efficiency

Comparison of different detectors concerning quantum efficiency between 200 – 1100 nm



Detector & Quantum Efficiency

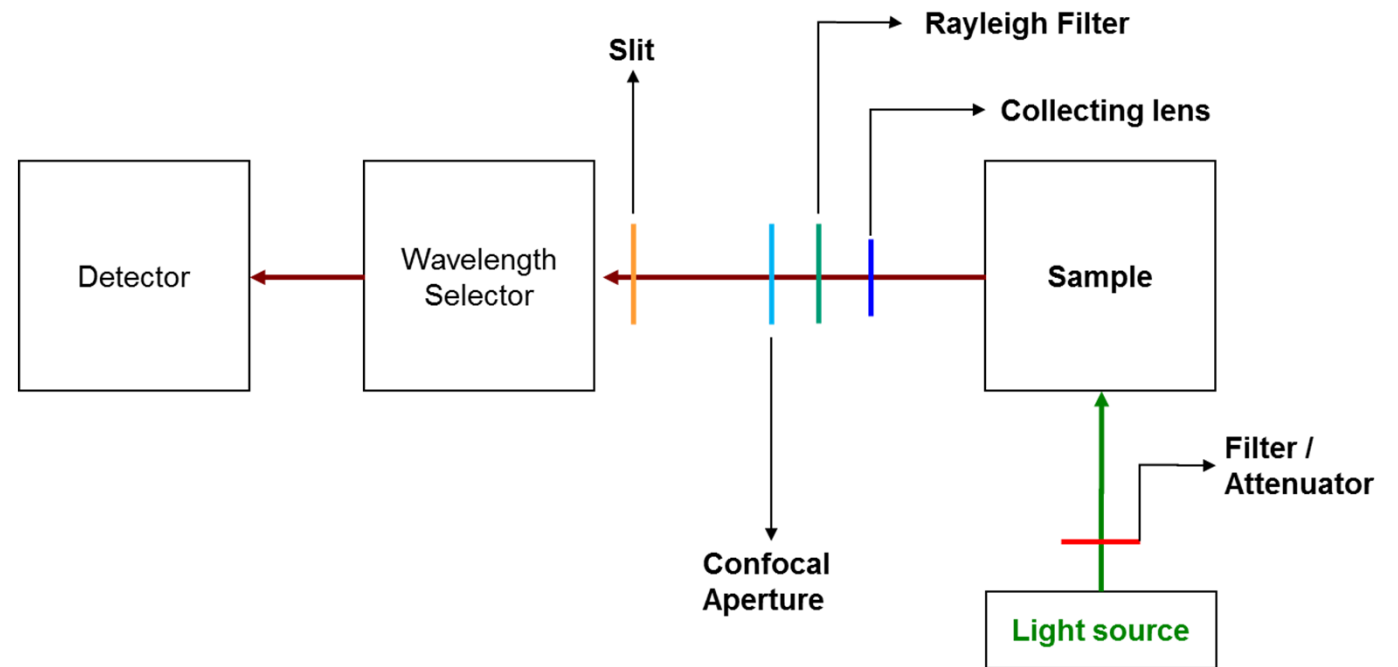


Outline

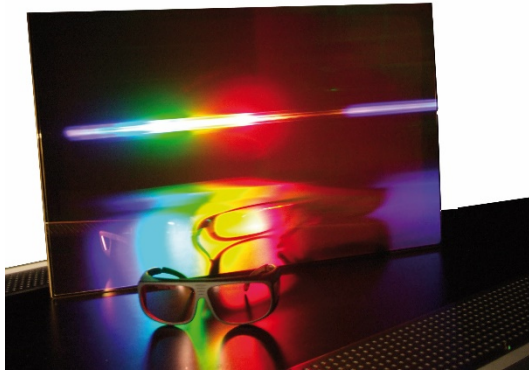
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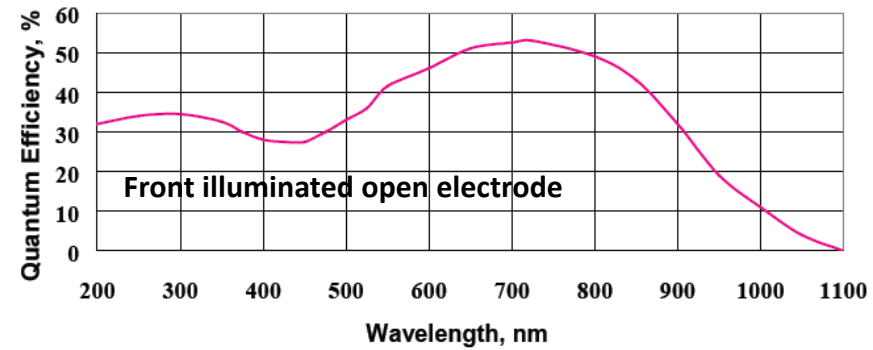
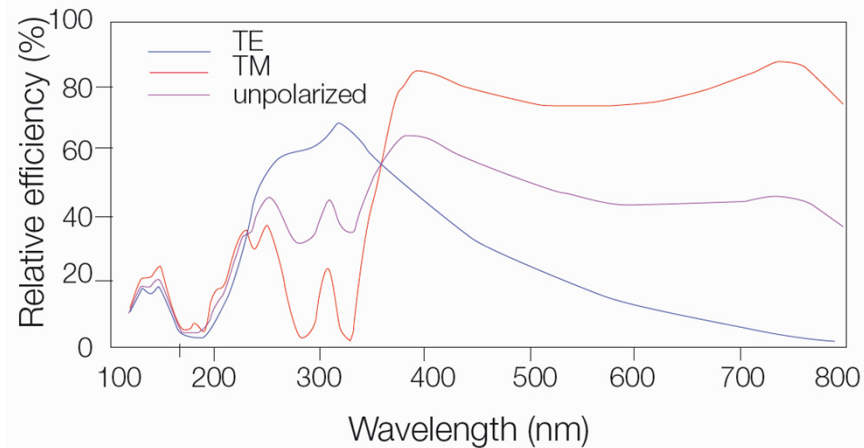


Correction of Instrument Function

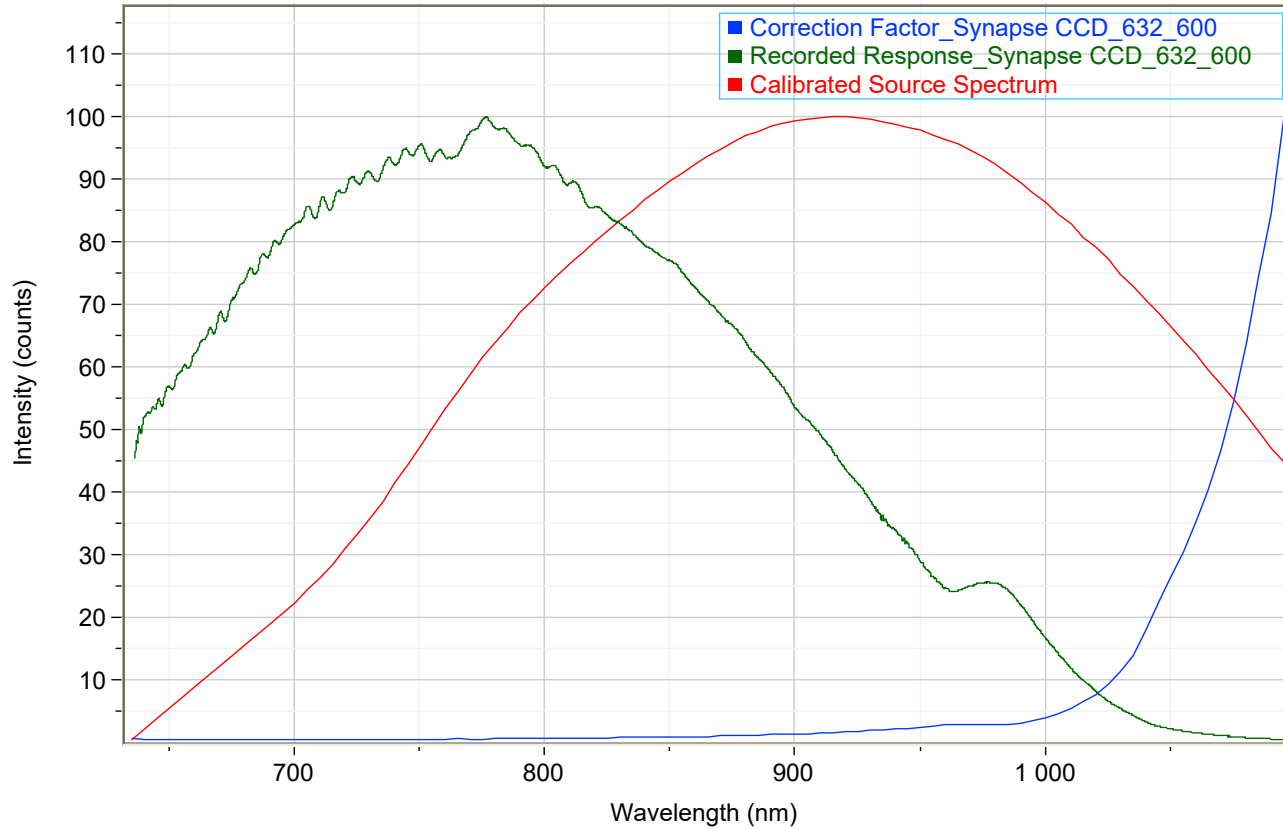


Reflectivity of the gratings vary with the wavelength as well as the transmission capability of lenses.

All this together (detector, gratings, mirrors, lenses) are generating a wavelength depended instrument function.



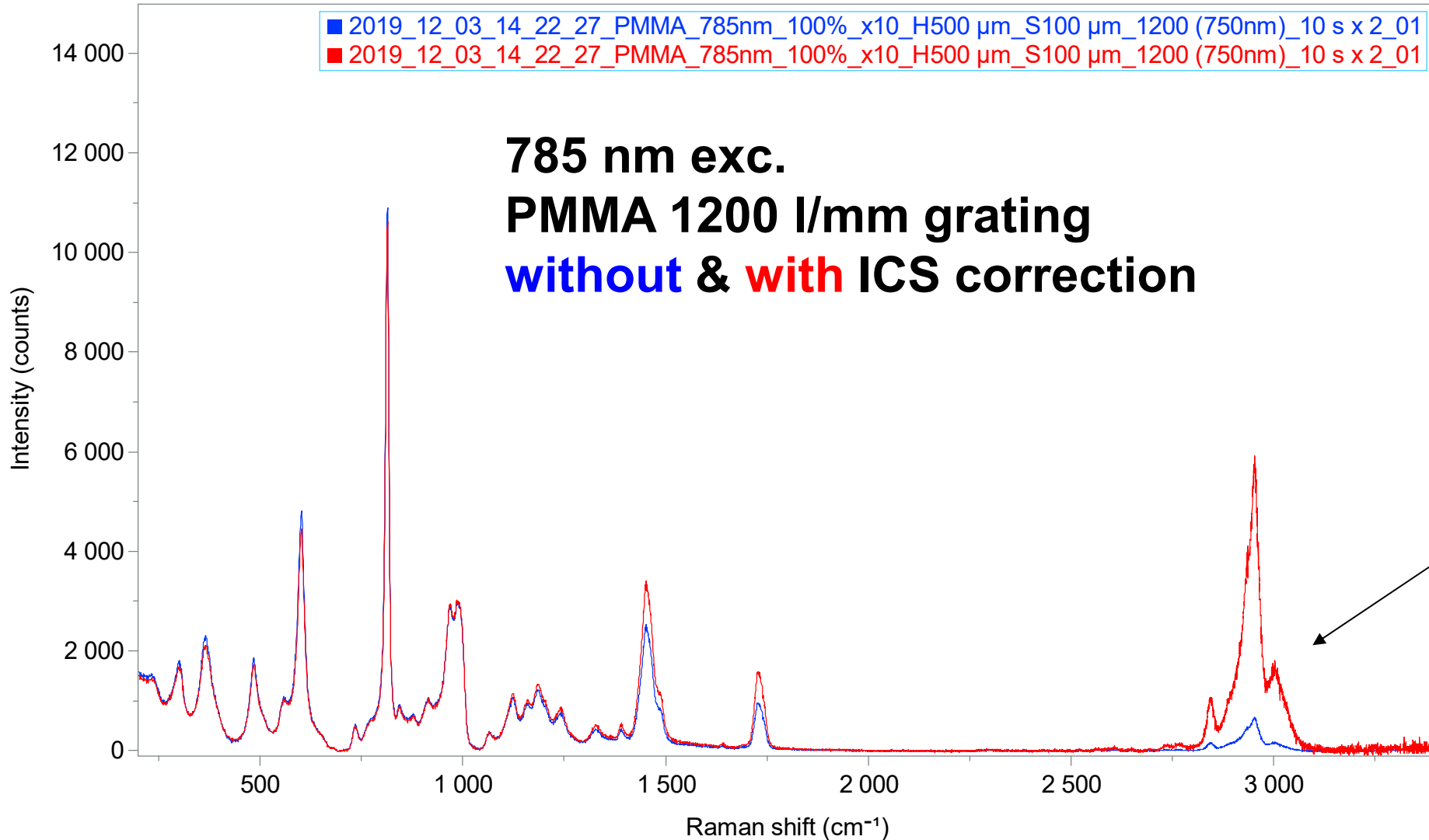
Correction of Instrument Function



▲ Acquisition options

Delay time (s)	0
Binning	1
Readout mode	Signal
Shutter mode	Auto
Spike filter	Off
Denoiser	Off
Laser mode	Auto
Trigger	Internal
ICS correction	On
Dark correction	Off

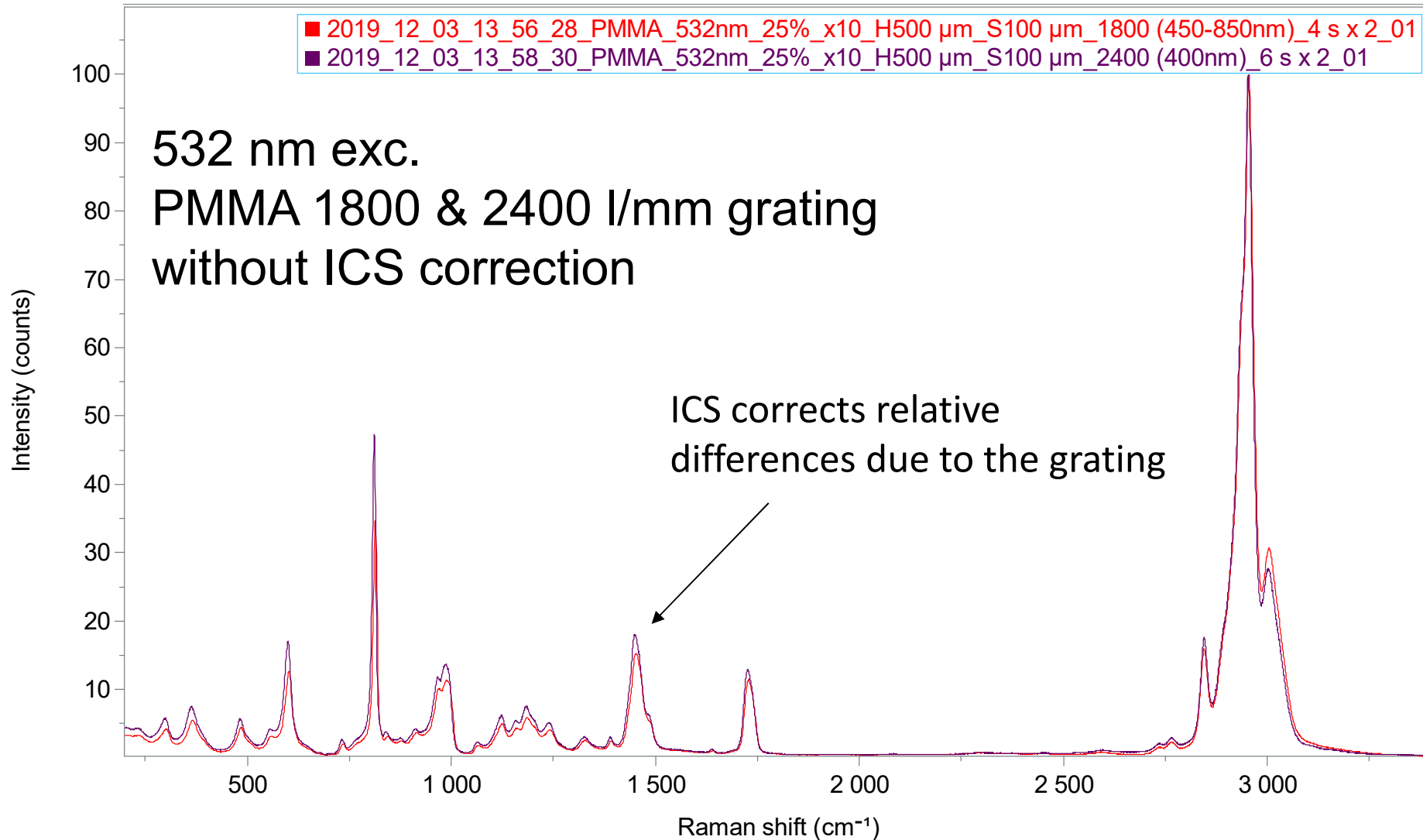
Correction of Instrument Function



Please consider: Correction of instrument function is a prerequisite for comparison of spectra and / or relative band intensities acquired with different conditions

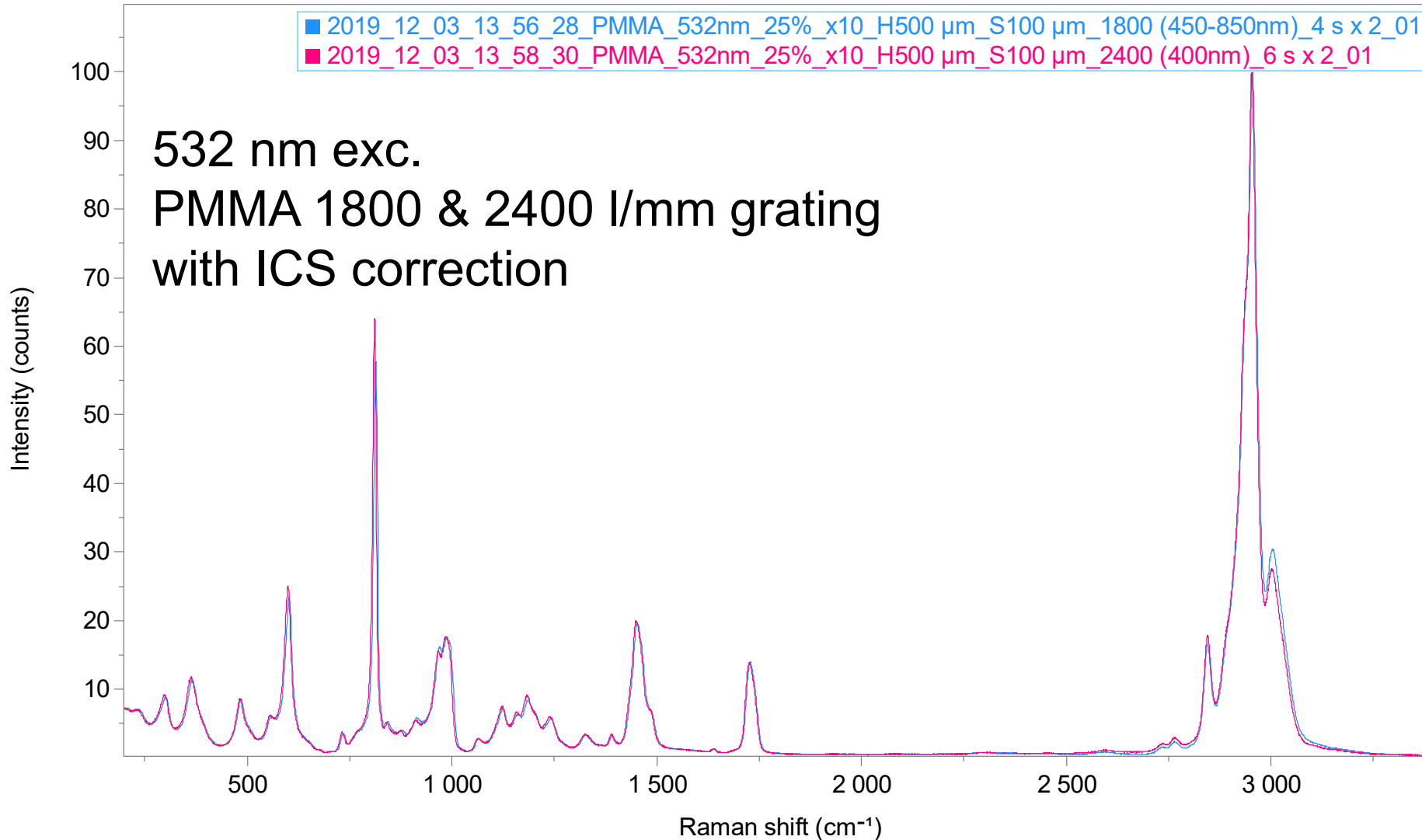
ICS corrects detector quantum efficiency

Correction of Instrument Function



Please consider: Correction of instrument function is a prerequisite for comparison of spectra and / or relative band intensities acquired with different conditions

Correction of Instrument Function



Please consider: Correction of instrument function is a prerequisite for comparison of spectra and / or relative band intensities acquired with different conditions

Thank you

Thank you

Omoshiro-okashiku
Joy and Fun



감사합니다 Cảm ơn

ありがとうございました

Dziękuję धन्यवाद Grazie

Merci 谢谢 நன்ற

ขอบคุณครับ Obrigado

Σας ευχαριστούμε

Tack ska ni ha

شُكْرًا

Большое спасибо

Danke

Gracias